

154

155 **Grade 6 - Middle School Discipline Specific Core Model**156 **Earth and Space Science**

157

158 From the introduction to the Middle School Earth and Space Sciences Standards in the  
159 NGSS:

160 Students in middle school develop understanding of a wide range of topics in  
161 Earth and space science (ESS) that build upon science concepts from  
162 elementary school through more advanced content, practice, and crosscutting  
163 themes. There are six ESS standard topics in middle school: 1) Space Systems,  
164 2) History of Earth, 3) Earth's Interior Systems, 4) Earth's Surface Systems, 5)  
165 Weather and Climate, and 6) Human Impacts. The content of the performance  
166 expectations are based on current community-based geoscience literacy efforts  
167 such as the Earth Science Literacy Principles (Wysession et al. 2012), and is  
168 presented with a greater emphasis on an Earth Systems Science approach. The  
169 performance expectations strongly reflect the many societally relevant aspects of  
170 ESS (resources, hazards, environmental impacts) as well as related connections  
171 to engineering and technology. While the performance expectations shown in  
172 middle school ESS couple particular practices with specific disciplinary core  
173 ideas, instructional decisions should include use of many practices that lead to  
174 the performance expectations. (NGSS Lead States 2013b)

175

176 A major emphasis of this course is the principle of interacting components of  
177 Earth's systems. As noted in the NRC Framework, "The natural and designed world is  
178 complex; it is too large and complicated to investigate and comprehend all at once.  
179 Scientists and students learn to define small portions for the convenience of  
180 investigation. The instructional segments of **investigations** can be referred to as  
181 '**systems**.' A system is an organized group of related objects or components that form a  
182 whole. Systems can consist, for example, of organisms, machines, fundamental  
183 particles, galaxies, ideas, and numbers. Systems have boundaries, components,  
184 resources, flow, and feedback (National Research Council [NRC] 2012)."

185

186 Although any real **system** smaller than the entire Universe interacts with and is  
187 dependent on other (external) systems, it is often useful to conceptually isolate a single  
system for study. To do this, scientists and engineers imagine an artificial boundary

188 between the system in question and everything else. Then they examine the system in  
189 detail while treating the effects of things outside the boundary as either forces acting on  
190 the system or flows of matter and energy across it—for example, the gravitational force  
191 due to Earth on a book lying on a table or the carbon dioxide expelled by an organism.  
192 Consideration of flows into and out of the system is a crucial element of system design.  
193 In the laboratory or even in field research, the extent to which a system under study can  
194 be physically isolated or external conditions controlled is an important element of the  
195 design of an investigation and interpretation of results.

196 Often, the parts of a **system** are interdependent – each one depends on or  
197 supports the functioning of the system’s other parts. Yet the properties and behavior of  
198 the whole system can be very different from those of any of its parts, and large systems  
199 may have emergent properties, such as the shape of a tree, that cannot be predicted in  
200 detail from knowledge about the components and their interactions. Things viewed as  
201 subsystems at one **scale** may themselves be viewed as whole systems at a smaller  
202 scale. For example, the circulatory system can be seen as an entity in itself or as a  
203 subsystem of the entire human body; a molecule can be studied as a stable  
204 configuration of atoms but also as a subsystem of a cell or a gas.

205 An explicit model of a system under study can be a useful tool not only for  
206 gaining understanding of the system but also for conveying it to others. **Models of a**  
207 **system** can range in complexity from lists and simple sketches to detailed computer  
208 simulations or functioning prototypes.

209 The systems identified in the Earth and space sciences course focus are the following:

- 210 • Atmosphere: gases around the Earth (i.e., our air)
- 211 • Hydrosphere: all the water (sometimes ice is separated out into the cryosphere).
- 212 • Geosphere: inorganic rocks and minerals
- 213 • Biosphere: all life
- 214 • Anthroposphere: humanity and all of its creations (primarily part of the biosphere,  
215 but often separated out to emphasize the significant influences humans have on  
216 the rest of Earth’s systems).

217  
218 Table 1 provides a schematic organization of the instructional segments and the primary  
219 Earth systems discussed in each instructional segment. The CA NGSS has titled this

220 discipline Earth and Space Sciences to emphasize that while Earth exists as a singular  
221 planet, its systems are strongly influenced by interactions with the broader Universe.

222 Table 1. Illustration of how different instructional segments relate to Earth's systems.

Instructional segment	Atmos.	Hydro.	Geo.	Bio.	Anthro.
Instructional segment 1: Earth's Place in the Solar System			X		
Instructional segment 2: Atmosphere: Cycles of Energy	X	X			X
Instructional segment 3: Atmosphere/Hydrosphere: Cycles of Matter	X	X			X
Instructional segment 4: Geosphere, External Processes		X	X	X	X
Instructional segment 5: Geosphere: Internal Processes			X		X

223

224 Each of these **systems** has components that interact with each other. Modeling  
225 the appropriate relationship between these components is at the center of each  
226 instructional segment in this course. Further, each system interacts with the others,  
227 originating the processes that shape our Earth.

228 In grade six, students apply and expand their prior understanding of these  
229 systems from their science experiences in fifth grade. Thus, beside grade-appropriate  
230 proficiency in using all the science and engineering practices and crosscutting concepts,  
231 students have developed an understanding of Earth's major systems (5-ESS2-1;  
232 ESS2.A) aided by concepts in physical science (PS1: structure and properties of matter;  
233 PS3.D: energy in chemical processes and everyday life) and life science (LS2.B: Cycles  
234 of matter and energy transfer in ecosystems). Table 2 shows the disciplinary core ideas

235 that students in sixth grade have experienced in fifth grade or earlier grades. Sixth  
 236 grade teachers will have to probe using a variety of formative and diagnostic  
 237 assessment tools the level of familiarity and mastery that their students have as they  
 238 enter their sixth grade science classes.

239 Table 2: Disciplinary Core Ideas included in grade 5. Shaded in gray are the core ideas  
 240 that pertain to Earth and Space Sciences.

Disciplinary Core Ideas	PS1: Matter and Its Interactions	PS1.A: Structure and properties of matter PS1.B: Chemical reactions
	PS2: Forces and Interactions	PS2.B: Types of interactions (gravitational force)
	PS3: Energy	PS3.D: Energy in chemical processes and everyday life
	PS4: Waves and Electromagnetic Interactions	(Addressed in grade 4)
	LS1: From Molecules and Organisms: Structure and Processes	LS1.C: Organization of matter and energy flow in organisms
	LS2: Ecosystems: Interactions, Energy, and Dynamics	LS2.A: Interdependent relationships in ecosystems LS2.B: Cycles of matter and energy transfer in ecosystems
	LS3: Heredity: Inheritance and Variation of Traits	(Addressed in grade 3)
	LS4: Biological Evolution: Unity and Diversity	(Addressed in grade 3)
	ESS1: Earth’s Place	ESS1.A: the Universe and its stars

	in the Universe	ESS1.B: Earth and the solar system
	ESS2: Earth's Systems	ESS2.A: Earth materials and systems ESS2.C: the role of water in Earth's surface processes
	ESS3: Earth and Human Activity	ESS3.C: Human Impacts on Earth systems

241

242 Earth and space sciences have much in common with the other branches of  
 243 science, but they also include a unique set of scientific pursuits. Inquiries into the  
 244 physical sciences (PS2) (e.g., forces, energy, gravity, magnetism) were conducted in  
 245 part as a means of understanding the size, age, structure, composition, and behavior of  
 246 Earth, the Sun, and the Moon; physics and chemistry later developed as separate  
 247 disciplines. The life sciences likewise are partially rooted in earth science, as Earth  
 248 remains the only example of a biologically active planet, and the fossils found in the  
 249 geological record of rocks are of interest to both life scientists and earth scientists (LS4).  
 250 As a result, the majority of research in Earth and space sciences is inter-disciplinary in  
 251 nature and is often organized into the categories of astrophysics, geophysics,  
 252 geochemistry, and geobiology. However, the underlying traditional discipline of geology,  
 253 involving the mapping and interpretation of rocks, remains a cornerstone of Earth and  
 254 space sciences.

255 When adapting the CA NGSS to their classroom, teachers have great  
 256 opportunities to make the subject matter regionally relevant. Coastal communities may  
 257 wish to focus on different spheres of interaction than farming communities in California's  
 258 Central Valley. Despite these regional differences, large portions of California's students  
 259 live in densely urban communities where ties to the natural environment are less  
 260 apparent. When describing possible directions for meeting the PE's, the framework  
 261 makes efforts to identify directions that will be most relevant for urban youth and  
 262 mentions specific activities relevant to urban geoscience.

263 **Example Course Mapping for an Earth and Space Science Course**

264 In this section, two types of tables have been included to provide an overview of the  
 265 materials contained:

- 266 1. A summary table: this table provides an overview of the suggested instructional  
 267 segments identified for this grade level.
- 268 2. Separate Instructional segment tables: these tables provide further details of the  
 269 three dimensions of the CA NGSS included in each instructional segment.

270

271 **Summary table for an example course in Middle School Earth and Space**  
 272 **Sciences.**

Instructional segment 1: Earth's Place in the Solar System	Performance Expectations addressed		
	MS-ESS1-1 MS-ESS1-2, MS-ESS1-3		
	Highlighted SEP	Highlighted DCI	Highlighted CCC
	<ul style="list-style-type: none"> <li>• Developing and using models</li> <li>• Analyzing and interpreting data</li> </ul>	ESS1.A: The Universe and its stars  ESS1.B: Earth and the solar system  <i>Other Necessary DCI:</i> PS2.B: Types of Interactions	<ul style="list-style-type: none"> <li>• Patterns</li> <li>• Scale, proportion, and quantity</li> <li>• Systems and system models</li> </ul>
	Summary of DCI		
	Patterns of the apparent motion of the sun, the Moon, and stars in the sky can be observed, described, predicted, and explained with models. Galaxies consist of stars, gases, and a collection of objects, including planets, their Moons, and asteroids that are held in orbit by gravitational forces.		

273

segment 2: Atmosphere: Cycles	Performance Expectations Addressed		
	MS-ESS1-1, MS-ESS2-6, MS-ESS3-4, MS-ESS3-5		
	Highlighted SEP	Highlighted DCI	Highlighted CCC
	<ul style="list-style-type: none"> <li>• Developing and using models</li> <li>• Planning and carrying out</li> </ul>	ESS2.C: The role of water in Earth's surface processes	<ul style="list-style-type: none"> <li>• Cause and effect</li> <li>• Systems and system models</li> </ul>

	<ul style="list-style-type: none"> <li>investigations</li> <li>Asking questions and defining problems</li> <li>Analyzing and interpreting data</li> </ul>	ESS2.D: Weather and climate ESS3.B: Natural hazards ESS3.D: Global climate change  <i>Other necessary DCI:</i> PS3.B: Conservation of Energy and Energy Transfer PS4.B: Electromagnetic radiation	<ul style="list-style-type: none"> <li>Patterns</li> </ul>
Summary of DCI			
Human activities, such as the release of greenhouse gases from burning fossil fuels, are major factors in the current rise in Earth’s mean surface temperature (global warming).			
Instructional segment 3: Atmosphere/Hydrosphere: Cycles of Matter	Performance Expectations Addressed		
	MS-ESS2-4, MS-ESS2-5, MS-ESS2-6, MS-ESS3-3, MS-ESS3-5		
	Highlighted SEP	Highlighted DCI	Highlighted CCC
	<ul style="list-style-type: none"> <li>Developing and using models</li> <li>Planning and carrying out investigations</li> <li>Asking questions and defining problems</li> <li>Analyzing and interpreting data</li> </ul>	ESS2.C: The role of water in Earth’s surface processes ESS2.D: Weather and climate ESS3.B: Natural hazards ESS3.D: Global climate change  <i>Other necessary DCI:</i> PS3.B: Conservation of Energy and Energy Transfer	<ul style="list-style-type: none"> <li>Cause and effect</li> <li>Systems and system models</li> <li>Patterns</li> </ul>
	Summary of DCI		
	Water continually cycles among land, ocean, and atmosphere via transpiration, evaporation, condensation and crystallization, and precipitation as well as downhill flows on land.		

274  
275  
276  
277

278  
279  
280  
281

Instructional segment 4: Geosphere, External Processes	Performance Expectations Addressed		
	MS-ESS1-4, MS-ESS2-1, MS-ESS2-2, MS-ESS3-1, MS-ESS3-2		
	Highlighted SEP	Highlighted DCI	Highlighted CCC
	<ul style="list-style-type: none"> <li>Analyzing and interpreting data</li> <li>Constructing explanations and designing solutions</li> </ul>	ESS1.C: The history of planet Earth ESS2.A: Earth’s materials and systems ESS2.C: The roles of water in Earth’s surface processes  Other necessary DCI: LS4.A: Evidence of Common Ancestry and Diversity	<ul style="list-style-type: none"> <li>Patterns</li> <li>Scale, proportion, &amp; quantity</li> </ul>
	Summary of DCI		
	The geological time scale interpreted from rock strata provides a way to organize Earth’s history.		
Instructional segment 5: Geosphere: Internal Processes	Performance Expectations Addressed		
	MS-ESS2-1, MS-ESS2-2, MS-ESS2-3, MS-ESS3-1, MS-ESS3-2		
	Highlighted SEP	Highlighted DCI	Highlighted CCC
	<ul style="list-style-type: none"> <li>Analyzing and interpreting data</li> <li>Constructing explanations and designing solutions</li> </ul>	ESS1.C: The history of planet Earth ESS2.A: Earth’s materials and systems ESS2.B: Plate tectonics and large-scale system interactions  Other necessary DCI: LS4.A: Evidence of Common Ancestry and Diversity	<ul style="list-style-type: none"> <li>Patterns</li> <li>Scale, proportion, &amp; quantity</li> </ul>
	Summary of DCI		



	Plate tectonics is the unifying theory that explains the past and current movements of the rocks at Earth’s surface and provides a framework for understanding the geological history.
--	--

282

283

284 **Grade 6 Instructional segment 1: Earth’s Place in the Solar System**

<b>Instructional segment 1: Earth’s Place in the Solar System</b>
<p>Guiding Questions:</p> <ul style="list-style-type: none"> <li>• What causes the patterns and cycles of stars, planets, and the moon?</li> <li>• How can we represent the vastness of the solar system and compare objects as large as planets and moons?</li> </ul>
<p>Highlighted Scientific and Engineering Practices:</p> <ul style="list-style-type: none"> <li>• Developing and using models</li> <li>• Analyzing and interpreting data</li> </ul>
<p>Highlighted Cross-cutting concepts:</p> <ul style="list-style-type: none"> <li>• Scale, proportion and quantity</li> <li>• Cause and Effect</li> </ul>
<p>Students who demonstrate understanding can:</p> <p><b>MS-ESS1-1. Develop and use a model of the Earth-Sun-Moon system to describe the cyclic patterns of lunar phases, eclipses of the Sun and Moon, and seasons.</b> [Clarification Statement: Examples of models can be physical, graphical, or conceptual.]</p> <p><b>MS-ESS1-2. Develop and use a model to describe the role of gravity in the motions within galaxies and the solar system.</b> [Clarification Statement: Emphasis for the model is on gravity as the force that holds together the solar system and Milky Way galaxy and controls orbital motions within them. Examples of models can be physical (such as the analogy of distance along a football field or computer visualizations of elliptical orbits) or conceptual (such as mathematical proportions relative to the size of familiar objects such as their school or state).] [Assessment Boundary: Assessment does not include Kepler’s Laws of orbital motion or the apparent retrograde motion of the planets as viewed from Earth.]</p> <p><b>MS-ESS1-3. Analyze and interpret data to determine scale properties of objects in the solar system.</b> [Clarification Statement: Emphasis is on the analysis of data from Earth-based instruments, space-based telescopes,</p>

and spacecraft to determine similarities and differences among solar system objects. Examples of scale properties include the sizes of an object's layers (such as crust and atmosphere), surface features (such as volcanoes), and orbital radius. Examples of data include statistical information, drawings and photographs, and models.] [Assessment Boundary: Assessment does not include recalling facts about properties of the planets and other solar system bodies.]

Significant Connections to California's Environmental Principles and Concepts:

None

## 285 **Background and instructional Suggestions**

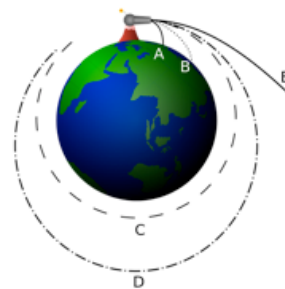
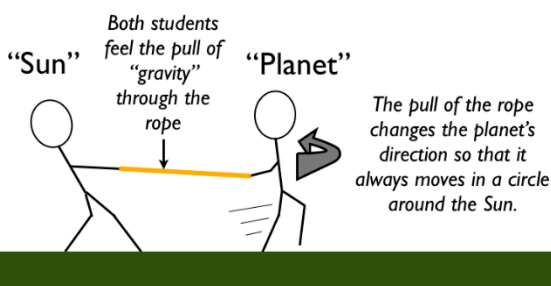
286 People throughout history have been fascinated by the heavens. Each ancient  
287 civilization noticed **patterns** in the movement of the Sun, Moon, and stars. Students  
288 themselves have recognized and described patterns of motion in the sky in earlier  
289 grades (**1-ESS1-1**, **5-ESS1-2**). Teachers can begin by reviewing those patterns,  
290 perhaps working with ELA teachers to read stories about the way in which ancient  
291 civilizations used the patterns in the stars to predict their motion. In this instructional  
292 segment, students construct **models** that explain the size, shape, and timing of these  
293 motions. Students should, like the people from ancient civilizations, be able to apply  
294 these models to qualitatively predict the motions of objects. In high school, they will  
295 extend this model by adding quantitative descriptions of the forces that cause the  
296 motion. Sixth grade lays the crucial foundation for that work.

297 Gravity is the driving force that shapes most of the motion in the Universe. In  
298 third grade, students investigated gravity as a force that can pull objects downward (3-  
299 PS2-1). If that's the case, why doesn't the moon fall down?<sup>1</sup> How can a force that pulls  
300 an object downward give rise to the ordered **patterns** we see in the movement of the  
301 stars in the sky? In this instructional segment, students **develop a model** of this  
302 process (**MS-ESS1-2**). Essential components of the model are 1) gravity is a force that  
303 pulls massive objects towards one another; 2) objects in the solar system move in

---

<sup>1</sup> NASA Ask an Astronomer, Why doesn't the Moon fall down,  
<http://www.spitzer.caltech.edu/video-audio/62-ask2002-001-Why-Doesn-t-the-Moon-Fall-Down->

304 circular patterns around the Sun and stars in galaxies move in circular patterns around  
 305 their center. Students can illustrate the relationship between these ideas with a rope  
 306 (left side of Figure 1). One person stands in the center and holds the rope while the  
 307 other starts moving away. Once the rope is taught, both people feel the rope tugging  
 308 them together. The pull of the rope changes the moving person's direction, constantly  
 309 pulling that person back on course so that he or she moves only in a circular motion.  
 310 Isaac Newton developed a conceptual model of this with the idea of a cannon shot from  
 311 a tall mountain at different speeds. Gravity always pulls the cannon ball down, but the  
 312 direction of "down" changes constantly (just like the *direction* of pull from the rope  
 313 changes constantly as the student runs around the circle). Online interactive simulations  
 314 of Newton's cannon can help students visualize the model even better.



315

316 Figure 1. Models showing the relationship between gravity and the circular motion of  
 317 objects in orbits. The left side is a physical model with students representing planets.  
 318 The right side shows Newton's cannon, a conceptual model illustrated in a diagram.  
 319 Image Credit: (CC-BY-NC-SA) by Matthew d'Alessio (LEFT) and Brondel 2010

320 The clarification statement for *MS-ESS1-2* may cause confusion, since many of  
 321 the examples given pertain to **scale models** for *MS-ESS1-3*. The two PE's are  
 322 intricately connected because gravity and motion help define the shape and scale of  
 323 recognizable bodies in our solar system. The next section describes some of these  
 324 relationships.

### 325 **Common Core Connection: Solar system scale**

326 When pondering Earth's place within the solar system, **scale and proportion** are  
 327 repeating concepts, and they align well with the **mathematical thinking** about ratios  
 328 and proportions from 6<sup>th</sup> grade mathematics (*CA CCSSM6.RP.1*). NASA has a series of

329 activities on Solar System math that allow students **analyze data** about solar system  
330 **scale** and then build scale **models**<sup>2</sup>. The activity from Day 1 in the vignette below is a  
331 related example pertaining to the Moon. Students also can get a tangible sense of the  
332 relative scale of the solar system by constructing a scale model on a 100 yard football  
333 field. Most of these examples provide solar system sizes as numbers in tables, but the  
334 clarification statement for *MS-ESS1-3* identifies another of other ways that students can  
335 obtain their **data for analysis**, including photographs, drawings, and models. For  
336 example, students can use online interactive models of the solar system to record the  
337 orbital distance and period of different planets. As the distance from the Sun increases,  
338 the time it takes for the planet to complete one orbit also increases. A similar activity  
339 can be done using a virtual telescope to analyze the orbital distance and orbital period  
340 of the moons of Jupiter<sup>3</sup>. A motivation for choosing to investigate orbital periods and  
341 radii is that it prepares students for calculating orbital periods using Kepler's Laws in  
342 high school (*HS-ESS1-4*).

343

#### 344 **Patterns in the Sun-Earth-Moon system**

345 The study of the moon under the *CA NGSS* illustrates some of the shifts in  
346 expectations compared to the 1998 California standards. Under the 1998 standards, 3<sup>rd</sup>  
347 grade students should **know** “the way in which the Moon’s appearance **changes** during  
348 the four week lunar cycle.” In the *CA NGSS*, students use observations to describe  
349 **patterns** in the moon’s motion in 1<sup>st</sup> grade (*1-ESS1-1*). Explaining the moon’s  
350 appearance is now part of 6<sup>th</sup> grade, but the emphasis is on **developing a model** that  
351 students can use to make and test predictions instead of simply describing the phases  
352 (*MS-ESS1-1*). The vignette below illustrates a teaching sequence that helps accomplish  
353 this model development.

354

---

<sup>2</sup> NASA, Solar System Math: <http://quest.nasa.gov/vft/#wtd>

<sup>3</sup> An older software package is here: Project CLEA: <http://www3.gettysburg.edu/~marschal/clea/juplab.html>, but a more modern version could be produced.

355

356

**Middle School Vignette**

357

**Using Models of Space Systems to Describe and Explain Patterns of Moon's**

358

**Phases**

359

(Adapted from NGSS Lead States 2013a, Case Study 3)

360

The vignette presents an example of how teaching and learning may look like in

the classroom when the *CA NGSS* are implemented. The purpose is to illustrate how a

teacher engages students in three-dimensional learning by providing them with

experiences and opportunity to develop and use the science and engineering practices

and the crosscutting concepts to understand the disciplinary core ideas associated with

the topic in the instructional segment.

366

It is important to note that the vignette focuses on only a limited number of

performance expectations. It should not be viewed as showing all instruction necessary

to prepare students to fully achieve these performance expectations or complete the

instructional segment. Neither does it indicate that the performance expectations should

be taught one at a time.

371

The vignette uses specific classroom contexts and themes, but it is not meant to

imply that this is the only way or the best way in which students are able to achieve the

indicated performance expectations. Rather, the vignette highlights examples of

teaching strategies, organization of the lesson structure, and possible students'

responses. In addition, science instruction should take into account that student

understanding builds over time and that some topics or ideas require activating prior

knowledge and extend that knowledge by revisiting it throughout the course of a year.

378

**Introduction**

379

The students in Mr. O sixth grade classroom receive science instruction five days

a week for 50 minutes each day. The students also receive instruction in

reading/language arts and mathematics in an integrated fashion. Strategic grouping of

382 students provides opportunities for peer-to-peer collaboration, facilitating support for  
383 struggling students, including English Language Learner students.

384 In the lesson sequence in this vignette, Mr. O uses multiple means of  
385 representations for allowing students to make sense of the view of the Moon phases as  
386 seen from Earth. These representations include: computer models using planetarium  
387 software (available free online<sup>4</sup>), physical models (foam balls, a lamp, golf balls), and  
388 diagrams such as foldables (three-dimensional interactive graphic representations with  
389 templates available online). Engaging the students in these multiple experiences to  
390 explain the same phenomenon and allowing them to **develop their own models** or  
391 evaluate alternative representations of the same model facilitates students'  
392 development of a conceptual model of the Earth-Sun-Moon system. In addition, the  
393 multiple experiences support language development as students discuss and **ask**  
394 **questions** about the experiences.

395 Mr. O has been preparing for this instructional segment for the past four months  
396 and he had strategically alerted students to look at the Moon in the sky throughout  
397 multiple days and notice changes in what they saw. Also, he often started the day by  
398 showing pictures of the Moon he had taken with his cell phone or had found online and  
399 had posted those picture in a corner of the classroom with a label indicating date and  
400 time. Most of the students knew already that the Moon appears different across different  
401 days of the month. Most of them, however, had not observed the Moon during daytime  
402 and they were surprised when Mr. O pointed out to the Moon in the sky one morning  
403 while they were playing in the playground before class. This preparation increased  
404 students' interest and everybody was excited when Mr. O announced, "Today we are  
405 going to learn about the Moon."

#### 406 **Day 1 - Exploring the Earth-Moon-Sun relationship.**

407 Mr. O initiated the instructional segment by asking students to open their  
408 notebooks, write the numbers 1-8 down the next blank page, and title it "Relative

---

<sup>4</sup> Stellarium: <http://stellarium.org>

409 Diameters.” On the interactive whiteboard, he projected a slide from a multi-media  
410 presentation *Two Astronomy Games* that showed nine images each identified by a letter  
411 and a label (Morrow 2004). The images were the Sun, Earth, a space shuttle, the  
412 Moon, the solar system, Mars, a galaxy, and Jupiter. Students were asked to number  
413 the objects in order from smallest (number 1) to largest (number 8) and from nearest to  
414 the surface of the Earth to farthest from the surface of the Earth. As the students  
415 marked their choices on their own, Mr. O walked among the students’ tables to gain  
416 insight regarding their prior knowledge. He planned to have students come back to this  
417 page later. Kevin, one of the most talkative students, seemed pleased and announced,  
418 “I love to study space!”

419 Mr. O moved to the front of the classroom and picked up a standard-sized  
420 playground ball in his hand. He asked the class to imagine the ball was Earth and he  
421 wrote down the class’ consensus of the ball’s dimensions that they had measured in  
422 math class. The diameter of the ball was 42 cm. Then he presented the class with a box  
423 of seven balls in a variety of sizes and listed their dimensions on the interactive  
424 whiteboard. He asked: “If Earth was the size of this playground ball, which of these  
425 balls would be the size of the Moon?” One student (from each table) came up and  
426 chose the ball they thought would be correct. Their choices varied from a softball to a  
427 small marble.

428 Before going further, the class reviewed the term diameter and Mr. O asked, “If  
429 you know that Earth’s diameter is 12,756 kilometers and the Moon’s diameter is 3,476  
430 kilometers, with your table groups, come up with a method to see if the ball you chose is  
431 the right size for this size Earth (holding up the playground ball).” (**using mathematics  
432 and computational thinking**) (**scale, proportion, and quantity**) (CA CCSSM .6.RP.1)

433 After some discussion time, students reported their calculations. One group  
434 noticed that there was a proportional relationship in the diameters of approximately 1:4,  
435 Earth to Moon. A student asked how they made that determination. Jeff responded, “If  
436 you estimate using 12,000 and 3,000, three goes into twelve four times.” He showed on  
437 the interactive whiteboard how four circles of a Moon model fit across the diameter of an  
438 Earth model. Mr. O said, “Now look at your ball as a Moon model and decide if you

439 think it is the correct size. What can you do to be sure? Decide on a process.” He let  
440 them use the playground ball as needed. (*MS-ESS1.A*)

441 Each group reported their findings and methods for determining whether or not  
442 their choice would be correct. One group made lines on paper where the endpoint of  
443 their ball was and did the same for the playground ball. Using those measurements and  
444 the 1:4 ratio, they decided if their Moon was the correct size. Another group used string  
445 to measure the diameter of the balls and then determined whether or not it was correct.  
446 Still another group held their ball up against the playground ball and moved their ball  
447 four times while marking the playground ball with a finger to see if their ball was the  
448 correct size for the model of Earth.

449 All groups reported their findings to the classroom. Kevin was agitated as he  
450 explained, “I told my group they were not right. The racquetball is the only one that is  
451 possible as the Moon, but they wouldn’t believe me.” Mr. O asked Kevin to restate the  
452 rule for when his group disagrees. Kevin thought and said, “When my group disagrees, I  
453 listen and then tell them what I think.” The classroom came to a consensus that the  
454 racquetball was the correct size ball to represent the Moon for the playground ball to  
455 represent the Earth.

## 456 **Day 2 - Exploring the Earth-Moon-Sun relationship.**

457 On the next day, Mr. O showed the students the actual distance from Earth to the  
458 moon and the circumference of Earth in kilometers. He asked them to figure out the  
459 distance between Earth and Moon in the model and to show it using string. Students  
460 were shocked at the distance the Moon was from Earth in this model. Their estimates  
461 had been much lower.

462 The class continued this activity by choosing correct size balls for the sun and  
463 Earth. Students also considered the relative size of the Sun and the distance of the Sun  
464 from Earth in the model. They used the **evidence** of the diameter of the Sun and its  
465 distance from Earth in the same way they determined the size and distance of the Moon  
466 from Earth. Some students were surprised at the size of the Sun and its distance from  
467 Earth in this model. Jeff decided that they could not fit the Sun in the room. He



468 explained that it would take over 100 playground balls to approximate the Sun's  
469 diameter. Jeff was eager to share his mathematical skill at finding the answer: "I know  
470 the answer! It would take almost 12,000 playground balls lined up to show how far away  
471 the Sun would be in this model." (***scale, proportion, and quantity***)

472 The students returned to their initial ideas on the "Relative Diameters" page in  
473 their notebooks, renumber the objects, and write any ideas that had changed after  
474 making the model. After giving students time to record their responses, Mr. O showed  
475 images of the items on the interactive whiteboard and led a discussion of the great  
476 distances between objects in the solar system in preparation for modeling the Moon's  
477 phases. (MS-ESS1.B)

### 478 **Day 3 - Exploring Moon phases: Computer representation**

479 For this lesson sequence, Mr. O considered the make-up of the table groupings  
480 of students. He wanted all students to have support while determining methods to  
481 check their choice of the Moon model, so he grouped students with that concern in  
482 mind. He used physical representations of Earth and the Moon and had students  
483 represent the distance physically, thereby assisting them in visualization and  
484 comprehension.

485 Mr. O downloaded an open-source planetarium software onto his interactive  
486 whiteboard- connected computer as well as onto the 14 student computers he had in his  
487 classroom. Each student also received a one-page calendar and they were instructed  
488 to use it to collect data using the software. Mr. O launched the program on the  
489 interactive whiteboard, introduced the students to the software, and showed them how  
490 to change the date and set up the scale Moon so they could see the phases. Mr. O also  
491 showed how the Moon's and Earth's orbital planes are offset by 5 degrees in an effort to  
492 help students understand how light can illuminate the Moon when it is on the other side  
493 of Earth without being blocked by Earth's shadow.

494 Recording began on the first Sunday on the calendar and ended on the last  
495 Saturday, resulting in five weeks of **data to analyze (analyzing and interpreting data)**  
496 Mr. O modeled how to record the data on the whiteboard next to the interactive

497 whiteboard. Students recorded the time and direction of moonrise and moonset as well  
498 as the apparent shape of the Moon in the sky for each date. To make sure that  
499 students understood the process and were recording accurately, he walked through the  
500 room and checked student work throughout the lesson.

501 During this data collection process, the students were told to focus their attention  
502 to the Sun-Moon relationship so they could see the light from the Sun traveling in a  
503 straight line to the Moon. The Moon was in the sky as the Sun was rising, and they  
504 focused on the Moon so that they could use the model for predictions. Mr. O asked,  
505 “Does anyone know where the Sun is right now?” Brady responded, “It’s more to the  
506 east and still rising.” Using the time and date function in the program, Mr. O advanced  
507 the time to show the sunrise and said, “Look at the Sun and Moon. What pattern do you  
508 notice about the light on the Moon in relation to the Sun?” (***Patterns***) Hillary answered,  
509 “It is going from the Sun to the Moon.” Mr. O responded, “Hmm. The light travels in a  
510 straight path from the Sun to the Moon. You have already learned that light travels in a  
511 straight line. Can we use that information to predict the position of the Sun even if we  
512 can’t see it? Let’s try as we continue.” After collecting six days of data, Mr. O asked  
513 students to look at the pattern in their data and predict the time and direction for  
514 moonrise and moonset on the next day. Bringing their attention to the patterns in the  
515 data he asked, “What time do you think the Moon will set on this day? The last time  
516 was 12:09.” Mark said, “I think 12:59.” Mr. O advanced the time in the software until the  
517 moonset – at 13:08. Jeff called out, “So it is setting about an hour later each time.” To  
518 reinforce the language Mr. O will use on many occasions throughout the instructional  
519 segment the following question, “What does that tell us about the planets and the  
520 Moon? They all move...” and students responded, “...in predictable patterns.”

521 A student said, “So let’s see if that ***pattern*** continues the whole month.” Once Mr.  
522 O was satisfied that the students had a foundation for data collection and that they were  
523 not just copying numbers from the software into their worksheet calendar, he told them  
524 to move to their computers in partners so they could work more independently to  
525 complete the data collection on the calendar. The students continued to record data  
526 about sunrise and moonrise until all the days in the handout calendar were filled.

527

528 **Day 4 – Exploring moon phases: Physical representation**

529 After students worked at the computers to complete  
 530 the calendar, Mr. O started a related activity in which they  
 531 modeled Moon phases using Styrofoam balls, their heads,  
 532 and a lamp with a bare bulb. In small groups students  
 533 stood in a circle around a lamp representing the Sun,  
 534 holding a Styrofoam ball on a stick representing the Moon.  
 535 They held the ball at arm’s length and rotated their bodies

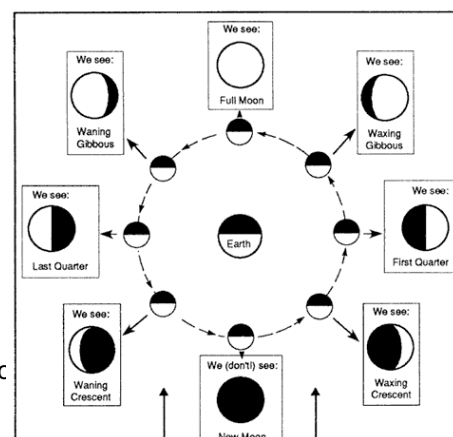


536 using their heads as a representation of Earth so they could see the earth view of the  
 537 Moon in all its phases in the lit portion of the ball. Mr. O directed Nicole to look at the  
 538 Styrofoam ball and the changing shadow. “What? I don’t see the shadow.” Mr. O  
 539 pointed out the curve of light on the Moon. “I see it!” Nicole said. The students went  
 540 through the phases, making a drawing in their notebook and naming each one. Small  
 541 groups allowed Mr. O to make sure that all students could see the lit portion on the  
 542 Styrofoam balls for each phase and were able to accurately illustrate the phases in the  
 543 model, giving him the opportunity to physically move them into position as necessary.  
 544 He frequently checked with students in the groups to show him to reproduce the  
 545 position of the Styrofoam ball corresponding to the drawings in their notebook.

546 For this activity, Mr. O expected all students to observe that the lit segment of  
 547 the Moon’s face increased, decreased, and increased again relative to the part in  
 548 shadow. He also expected students to notice that the lit side of the Moon was on the  
 549 left after the full Moon phase, and on the right after the new Moon phase, as viewed  
 550 from Earth.

551

552 **Day 5-7 – Developing a model to explain the**  
 553 **Moon phases.**



554 The next day, Mr. O pulled out large whiteboards and instructed the students to  
555 collaborate and make a drawing to explain how the model of the Moon phases  
556 illustrated changes in the apparent shape of the Moon. Mr. O started the lesson telling  
557 students they were going to make their thinking public by producing small group  
558 models. Students first organized their individual understanding by sketching and  
559 labeling the apparent changes in the moon based on what they observed and discussed  
560 in class. Next they took turns sharing their ideas with their group, noting similarities and  
561 differences. Mr. O walked the classroom listening to the progress as each group  
562 member shared. He reminded some groups of the classroom norms of respect and  
563 responsibility when participating in a group discussion. After reaching a consensus on  
564 the elements the group believed explained the apparent changes in the shape of the  
565 moon, they acquired a large whiteboard and produce a group consensus model to  
566 share with the class. They discussed limitations of the **models** – the things that a model  
567 is unable to show accurately. For example, the students identified the relative sizes of  
568 the Sun, Earth, and Moon as well as the relative distances between each as being  
569 inaccurate in this model. They had learned that in the previous days.

570 The following day, Mr. O announced they were doing a “Sticky Note” gallery walk  
571 of the models where each group would visit each of the models, consider and discuss  
572 them and then provide feedback on a color coded sticky note. Three different colors  
573 were used; one for questions, one for additions and one for suggested revisions.  
574 Students were reminded that the purpose of the feedback was to help the authors clarify  
575 the thinking that went into their model. Mr. O provided sentence frames to help students  
576 form questions, additions or suggested revisions. As the groups walked and discussed  
577 the use of the color-coding helped to focus their discussion and make it productive.

578 After completing the gallery walk, each group organized the sticky notes they  
579 received by the type of comment and then made revisions to their model based on the  
580 feedback.

581 Over the next two days Mr. O again pulled small groups of students to use  
582 another **physical model** showing Moon phases. This one used golf balls that were  
583 painted black on half of the sphere, leaving the other half showing the side of the Moon

584 lit by the Sun (Young and Guy 2008). The golf balls were drilled and mounted on tees  
 585 so they would stand up on a surface. Mr. O had two sets – one set up on a table that  
 586 showed the Moon in orbit around the earth in eight phase positions as the “space view”  
 587 model (Figure 1), and the other with the model Moons set on eight chairs circled in the  
 588 eight phase positions to show the “earth view” model (Figure 2).



Figure 1. Space view model



589 First, students were shown the space view model  
 590 and asked what they noticed about the Moons.  
 591 Mr. O wanted them to notice that the white sides of  
 592 all the balls (showing light) faced the same  
 593 direction. He asked them to identify the direction  
 594 of the Sun. Then Mr. O drew the students’  
 595 attention to the model on the chairs, the earth view  
 596 model. All the balls in this model faced the same  
 597 direction as those in the space view model.  
 598 Students again identified the direction of the Sun  
 599 and noted that the position of the Moons in both  
 600 **models** was the same (MS-ESS1.A). One at a  
 601 time, students physically got into the center of the  
 602 circle of chairs and viewed the phases at eye level,  
 603 which simulated the earth view of each phase.

604 Also, students compared their drawing on the whiteboard illustrating the model of the  
 605 Earth-Sun-Moon system with what they were seeing now. This activity made the  
 606 diagram, often found in books and worksheets showing both views on the same  
 607 diagram, less confusing to the students.

608 Throughout the lesson sequence, Mr. O continually formatively assessed  
 609 students’ progression of learning through observations and classroom discourse. If he  
 610 noticed students needed more experience with Moon phases, he provided them with  
 611 additional activities such as videos and Moon phase cards. In one formal assessment of  
 612 understanding, Mr. O paired students together so that one was assigned to be the earth  
 613 and the other the Moon. He designated one wall of the classroom as the Sun and then

614 asked the Moons to show different phases. The students switched roles so that Mr. O  
615 could assess everyone. He also used this model to demonstrate the Moon’s coincident  
616 rotation and revolution. In another formal assessment, he asked students to draw a  
617 model on whiteboards showing the relationship of the earth, Moon, and Sun in full Moon  
618 phase.

619 **Day 8 – Solidify learning about Moon phases and extend learning through**  
620 **readings.**

621 Mr. O brought all students together the next day to create a foldable showing the  
622 earth view of the Moon phases similar to diagrams found in books. Students created  
623 their Moon phases using eight black circles and four white circles, cutting the white  
624 circles to make two crescent Moons, two gibbous Moons and two-quarter Moons. The  
625 white circle pieces were placed on the black circles to create the phases, and later  
626 glued on the foldable.

627 Students partnered to read *The Moon* by Seymour Simon (2003). Students used  
628 the **information** in the book to label the Moon phases on their foldable, write about the  
629 Moon’s surface, and record any new **questions** that arose from their reading. Kevin  
630 asked, “When is the next solar and lunar eclipse?” Jeanette questioned, “What samples  
631 were brought back from the Moon?” And Nicole wanted to know, “Where did Americans  
632 land on the Moon?” To support their reading of the text, the teacher gave Hillary, Brady,  
633 and Jeff the option of being paired with students who had more advanced reading skills.  
634 The teacher allowed students who finished with the entire reading task to use text  
635 materials and Internet resources to research answers to the questions they developed  
636 when reading *The Moon*. The teacher will use these answers to these questions during  
637 the next few days. For example, Mr. O may have students revisit the physical model of  
638 the Earth-Sun-Moon system to explain solar and lunar eclipses (*MS-ESS1-1*).

639 **NGSS Connections and Three-Dimensional Learning**

**Performance Expectations**

**MS-ESS1-1 Earth’s Place in the Universe**

*Develop and use a model of the Earth-sun-moon system to predict and describe the cyclic patterns of lunar phases, eclipses of the sun and moon, and seasons.*

**MS-ESS1-3 Earth’s Place in the Universe**

*Analyze and interpret data to determine scale properties of objects in the solar system.*

<b>Science and engineering practices</b>	<b>Disciplinary core ideas</b>	<b>Cross cutting concepts</b>
<p><b>Developing and Using Models</b></p> <p><i>Develop and use a model to describe phenomena.</i></p> <p><b>Analyzing and Interpreting Data</b></p> <p><i>Analyze and interpret data to determine similarities and differences in findings</i></p>	<p><b>ESS1.A The Universe and Its Stars</b></p> <p><i>Patterns of the apparent motion of the sun, the moon, and stars in the sky can be observed, described, predicted, and explained with models.</i></p> <p><b>ESS1.B Earth and the Solar System</b></p> <p><i>The solar system consists of the sun and a collection of objects, including planets, their moons, and asteroids that are held in orbit around the sun by its gravitational pull on them. This model of the solar system can explain tides and eclipses of the sun and the moon.</i></p>	<p><b>Patterns</b></p> <p><i>Patterns can be used to identify cause-and-effect relationships</i></p> <p><b>Scale, Proportion, and Quantity</b></p> <p><i>Time, space, and energy phenomena can be observed at various scales using models to study systems that are too large or too small.</i></p>
<p><b>Connections to the CA CCSSM: MP. 1, MP. 2, MP. 3</b></p>		

**Connections to CA CCSS for ELA/Literacy:** RI.7, RI.8, SL.8.1, SL 8.2, RST.6-8.2, RST.6-8.3

**Connection to CA ELD Standards:** ELD.PI.8.1, ELD.PI.8.5, ELD.PI.8.6a-b,

640

### 641 **Vignette Debrief**

642       The *CA NGSS* require that students engage in science and engineering practices  
643 to develop deeper understanding of the disciplinary core ideas and crosscutting  
644 concepts. The lessons give students multiple opportunities to engage with the core  
645 ideas in space science (moon phases), helping them to move towards mastery of the  
646 three components described in the *CA NGSS* performance expectation.

647       In this vignette, the teacher selected two performance expectations and in the  
648 lessons described above and he engaged students only in selected portions of these  
649 PEs. Full mastery of the PEs will be achieved throughout subsequent instructional  
650 segments.

651       Students were engaged in a number of science practices with a focus on  
652 **developing and using models** and **analyzing and interpreting data**. Space science  
653 lends itself well to the use of models to describe **patterns** in phenomena and to  
654 construct explanations based on **evidence**.

655       With guidance from their teacher, students used the ratios of the diameters of  
656 Earth and its moon to construct a class model of the relative sizes of the two objects.  
657 Using distance and Earth's diameter or circumference ratios, they also constructed a  
658 distance model of those objects. In addition, the relative size of the Sun and the relative  
659 distance from Earth in this model was calculated and described, although not  
660 constructed (due to the constraints of the room and location). Throughout the vignette,  
661 a variety of **models** were used to help students identify **patterns** in the relative  
662 positions of the Earth, Moon and Sun, and to explain moon phases.



663 Students made predictions about the data collected and recorded them on the  
664 calendar, using the lens of the crosscutting concept of **patterns**. When analyzing and  
665 **interpreting the data**, they identified the patterns in the Earth-Moon-Sun relationship.  
666 The pattern made by the lit portion of the moon was observed and recorded. In  
667 addition, students considered the crosscutting concept of **scale, proportion, and**  
668 **quantity** as they constructed **models** of relative sizes and distance of the sun and  
669 planets.

### 670 **CCSS Connections to English Language Arts and Mathematics**

671 Students are engaged in small group work activities, both listening to their peers  
672 ideas and sharing their own thoughts. Students used the text in *The Moon Book* to label  
673 each phase of the Moon in their graphic organizer foldable. This connects to the CA  
674 *CCSS for ELA/Literacy* Reading Informational Text standard (RI.7). In addition, they  
675 summarized information about the surface of the moon inside their foldable, which  
676 corresponds to Reading Informational Text Standard 8 (RI.8).

677 When comparing sizes and distances, students were challenged to find ways of  
678 comparing numbers, applying the CA CCSSM Standard for Mathematical Practice 1  
679 (MP.1). In addition, students used rounding and estimation to calculate the quotients in  
680 the ratios, both skills developed in earlier grades. Throughout the instructional segment,  
681 students reasoned quantitatively as they compared the sizes of the Earth and Moon,  
682 Standard for Mathematical Practice 2 (MP.2). As students made conclusions about  
683 which ball was the moon, they argued for their selection and agreed or disagreed with  
684 each other using their calculation, Standard for Mathematical Practice 3 (MP.3)

685 **MP.1** Make sense of problems and persevere in solving them.

686 **MP.2** Reason abstractly and quantitatively.

687 **MP.3** Construct viable arguments and critique the reasoning of others.

688 **Resources for the Vignette**

689  
690  
691  
692  
693  
694  
695  
696

- Morrow, C. 2004. Two Astronomy Games.  
[http://www.spacescience.org/education/instructional\\_materials.html](http://www.spacescience.org/education/instructional_materials.html) (accessed August 5, 2015).
- Simon, S. 2003. *The Moon*. New York, NY: Simon and Schuster.
- Young, T., and M. Guy. 2008. "The Moon's Phases and the Self Shadow."  
*Science and Children* 46 (1): 30.

697

698

699 **Grade 6 Instructional segment 2: Atmosphere: Cycles of Energy**

Instructional segment 2: Atmosphere: Cycles of Energy
<p>Guiding Questions:</p> <ul style="list-style-type: none"> <li>• Why is it cold at the North Pole? (Or, why does Santa wear a big red suit?)</li> <li>• What causes California’s summers to be hot and dry? What causes the changes between summer and winter?</li> <li>• Why is there more rain in northern California than Southern California?</li> <li>• What effect do humans have on Earth’s climate?</li> </ul>
<p>Highlighted Scientific and Engineering Practices:</p> <ul style="list-style-type: none"> <li>• <b>Ask questions</b></li> <li>• <b>Develop and use models</b></li> </ul>
<p>Highlighted Cross-cutting concepts:</p> <ul style="list-style-type: none"> <li>• <b><i>Patterns</i></b></li> <li>• <b><i>Energy and matter flow</i></b></li> <li>• <b><i>Cause and effect</i></b></li> </ul>
<p>Students who demonstrate understanding can:</p> <p><b>MS-ESS1-1. Develop and use a model of the Earth-Sun-Moon system to describe the cyclic patterns of lunar phases, eclipses of the Sun and Moon, and seasons. [Clarification Statement: Examples of models can be physical, graphical, or conceptual.]</b> (Continued from instructional segment 1)</p> <p><b>MS-ESS2-6. Develop and use a model to describe how unequal heating and rotation of the Earth cause patterns of atmospheric and oceanic circulation that determine regional climates. [Clarification Statement: Emphasis is on how patterns vary by latitude, altitude, and geographic land distribution. Emphasis of atmospheric circulation is on the sunlight-driven latitudinal banding, the Coriolis effect, and resulting prevailing winds; emphasis of ocean circulation is on the transfer of heat by the global ocean convection cycle, which is constrained by the Coriolis effect and the outlines of continents. Examples of models can be diagrams, maps and globes, or digital representations.] [Assessment Boundary: Assessment does not include the dynamics of the Coriolis effect.]</b></p> <p><b>MS-ESS3-4. Construct an argument supported by evidence for how increases in</b></p>

**human population and per-capita consumption of natural resources impact Earth’s systems.** [Clarification Statement: Examples of evidence include grade-appropriate databases on human populations and the rates of consumption of food and natural resources (such as freshwater, mineral, and energy). Examples of impacts can include changes to the appearance, composition, and structure of Earth’s systems as well as the rates at which they change. The consequences of increases in human populations and consumption of natural resources are described by science, but science does not make the decisions for the actions society takes.]

**MS-ESS3-5. Ask questions to clarify evidence of the factors that have caused the rise in global temperatures over the past century.** [Clarification Statement: Examples of factors include human activities (such as fossil fuel combustion, cement production, and agricultural activity) and natural processes (such as changes in incoming solar radiation or volcanic activity). Examples of evidence can include tables, graphs, and maps of global and regional temperatures, atmospheric levels of gases such as carbon dioxide and methane, and the rates of human activities. Emphasis is on the major role that human activities play in causing the rise in global temperatures.]

---

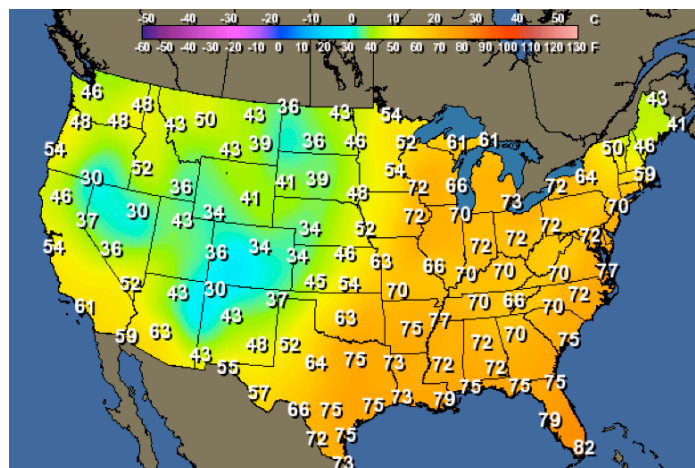
Significant Connections to California’s Environmental Principles and Concepts:

Principle III. Natural systems proceed through cycles that humans depend upon, benefit from and can alter.

Principle IV. The exchange of matter between natural systems and human societies affects the long term functioning of both.

700 **Background and instructional Suggestions**

701



702

703 **Figure 2.** A snapshot of morning temperatures across the USA reveals the importance  
704 of sunlight in affecting the temperature near Earth's surface. Image Credit:

705

706 During middle school, students identify some basic **patterns** in Earth's climate  
707 and **develop a model** of the factors that **cause** those patterns. The model is simple and  
708 related primarily to one part of Earth's **energy** balance, the input from the Sun. They  
709 extend this model in high school, so it is important to build this basic foundation (*HS-*  
710 *ESS2-4*).

711 Their **model** begins at the simplest level with recognizing that the more intense  
712 the solar input, the warmer temperatures are on Earth. Students can discover this  
713 pattern by looking at a map of temperature in the early morning across the USA. Look  
714 at **Figure 2** and draw a line dividing the country in half. What explains this simple  
715 **pattern**? The Sun has risen already across the east coast and has warmed it up.  
716 Students can also identify other trends such as the warming towards the southern half  
717 of the country and the behavior of California that appears warmer than its neighbors  
718 despite the fact that the Sun has not yet risen.

### 719 **Average temperature versus latitude**

720 **Figure 2** is just a snapshot in time that quickly changes, but there are also trends  
721 that last much longer. Should you bring beach clothes or a warm coat on a trip to  
722 Antarctica? How about San Diego, where it has only snowed 5 times there in the last  
723 125 years? How about Lake Tahoe, which typically receives more than 10 feet of snow  
724 in the winter but is a popular recreation area for swimmers and boaters every summer.  
725 Different cities tend to have predictable **patterns** in their weather that depend on the  
726 city's location and the time of year (their 'climate'). Students **investigate** these patterns

727 across the globe by **obtaining temperature information** from web sources<sup>5</sup> or from a  
728 simplified version for teaching<sup>6</sup>.

### 729 **Common Core Connection**

730 Students plot climatograms showing the average temperature for each month (CA  
731 CCSSM 6.SP.4). They calculate the average temperature of each city over the entire  
732 year, as well as its spread throughout the year (CA CCSSM 6.SP.2, CA CCSSM  
733 .6.SP.3).

734 Constructing a dot-plot with average annual temperature versus latitude and  
735 temperature spread versus latitude reveals an important **pattern**. Students probably  
736 already knew that it was cold at the North Pole, but why is there such a large  
737 temperature range at the poles and not at the equator? Within 20 degrees of the  
738 equator and 20 degrees of the poles, latitude doesn't have a major impact on climate  
739 and cities share fairly similar climates to one another. In between these sections of the  
740 Earth, climate varies greatly with latitude. Students should start **asking questions**  
741 about the cause of these differences.

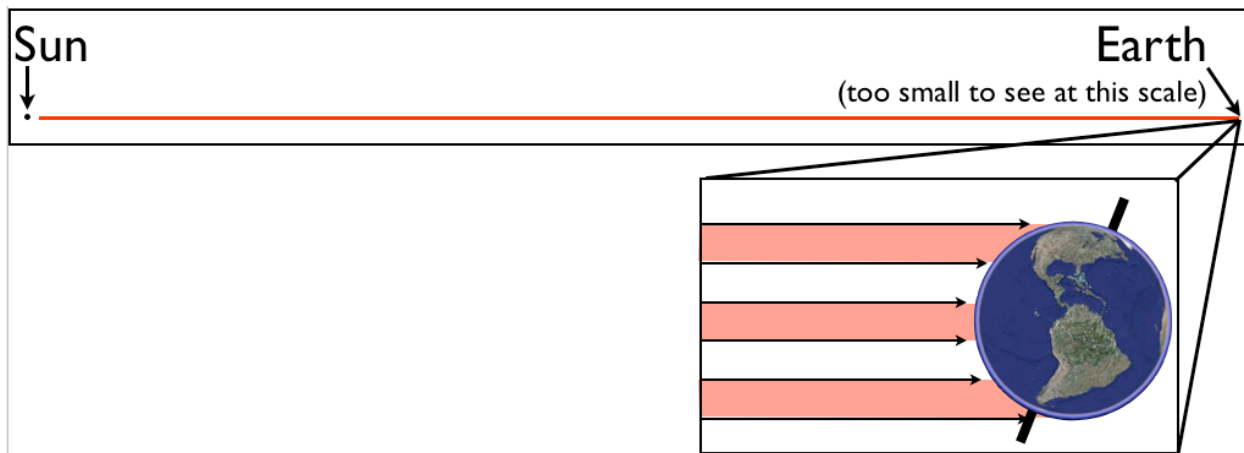
742 Like the temperature map in **Figure 2**, these long-term temperature differences relate to  
743 the difference in energy received from the Sun. How can the equator appear to receive  
744 more energy than the either of the poles despite the fact that they all receive their  
745 energy from the same Sun? The key is that the Earth is a sphere. Sunlight arrives at  
746 Earth as parallel rays, but hits the surface at nearly a 90° angle near the equator and at  
747 flatter/smaller angles near the poles because of Earth's round shape. The light spreads  
748 out over a larger area near the poles, meaning that each square foot patch of the  
749 surface receives a smaller **proportion** of the energy coming from the Sun than that  
750 same patch does at the equator (**Figure 3**), which causes the sunlight on that patch to  
751 be less intense. When the sun shines down at a 90° angle, a patch of land receives

---

<sup>5</sup> NOAA, *Global Historical Climatology Network-Monthly (GHCN-M)*:  
<http://www.ncdc.noaa.gov/ghcnm/v3.php>

<sup>6</sup> My version is at <http://zadok.org/climate>, but there is probably something better for middle school.

752 twice the energy compared to a 30° angle, so this effect has a big impact on the  
 753 temperature. So even though the entire Earth receives energy from the same Sun, each  
 754 section receives a different portion of the Sun's rays, depending on its latitude (**Figure**  
 755 **4**).



756

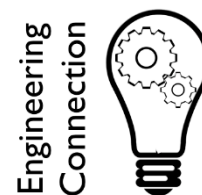
757 **Figure 3.** A scale illustration of the Earth-Sun system (top). The Sun is 5 pixels wide  
 758 and the Earth is 1075 pixels away, but is only 0.05 pixels wide, which is too small to  
 759 display. At this scale, it is easier to recognize that rays of sunlight arrive at Earth as  
 760 parallel rays at all latitudes (bottom). Image Credit: (CC-BY-NC-SA) M. d'Alessio.

761

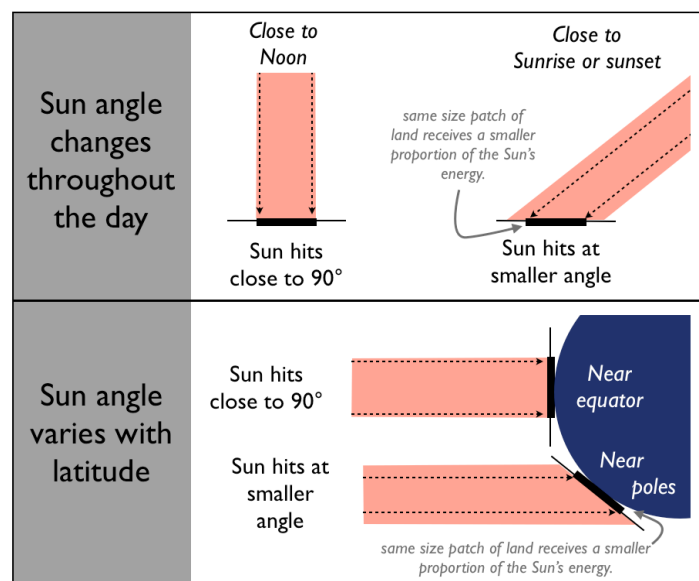
762 Students perform an **investigation** of the relationship between light intensity and  
 763 angle by shining a flashlight at a piece of paper at different angles while keeping the  
 764 distance between the light and the paper constant (NASA 2008). Students can directly  
 765 observe how the patch of light gets dimmer when it strikes the page at a low angle and  
 766 spreads out over a large area. While a piece of paper is flat, students simulate the  
 767 parallel rays of sunlight arriving at Earth by shining their flashlight on a round ball and  
 768 observing how the patch of light is small and intense near the equator but spread out  
 769 near the poles.

### 770 **Engineering connection: Solar array design**

771 This concept has important engineering applications for  
 772 solar energy. California hosts several of the world's largest  
 773 arrays of solar panels in the world. When people place solar



774 panels on their roofs, the angle of the panels is usually fixed by the angle of the roof. To  
 775 maximize efficiency at large solar power arrays, the motors constantly turn the panels  
 776 so that they face the Sun at an angle as close to  $90^\circ$  as possible to get the maximum  
 777 energy output. Students can experience this effect in a classroom with a small solar  
 778 panel hooked up to an electric motor. As they rotate the solar panel to change the angle  
 779 of sunlight, the energy output **changes** so that the motor turns at a different speed (New  
 780 York State Energy Research and Development Authority 2015). Students could engage  
 781 in an engineering challenge to design a rotating base for solar panels that has the  
 782 necessary range of movement (both tilting and swiveling) and uses low cost materials  
 783 (MS-ETS1-1, MS-ETS1-2).



784

785 **Figure 4.** Effect of the angle of the Sun's rays on area of the Earth's surface it  
 786 illuminates. At angles smaller than  $90^\circ$ , the energy is spread out over a larger area. The  
 787 effect is important as the sun moves across the sky during one day (top) and at different  
 788 latitudes across the planet (bottom). Image Credit: (CC-BY-NC-SA) M. d'Alessio.

789

## 790 **Uneven Heating and the Earth's 'Circulation System'**

791 The uneven heating between the equator and the poles is the root **cause** of all  
 792 earth's ocean and wind currents. They carry hot material (water in the oceans and air in  
 793 the atmosphere) from the equator towards the poles in a large-**scale** convection current.  
 794 Convection is a **cycling of matter** driven by the **flow of energy** (connects to MS-ESS2-



795 1, though assessment of that PE focuses largely on the solid Earth). As hot material  
796 moves poleward, colder material moves towards the equator. Without currents, regional  
797 temperatures would be extreme — super hot at the equator and frigid toward the  
798 poles—and much less of Earth’s land would be habitable. Sunlight heats Earth’s  
799 surface, which in turn heats the atmosphere. At the global scale, wind currents are  
800 dominated by three different directions of motion: 1) hot material rising vertically upward  
801 and cold material sinking vertically downward due to convection; 2) hot material from  
802 the equator moving northward towards the poles and cold material moving southward  
803 towards the equator due to convection; and 3) east-west apparent motion of material  
804 driven by Earth’s rotation. Ocean currents undergo similar motions modified by  
805 collisions with the coastlines that disrupt these ideal motions. While wind directions also  
806 change when they rise up over or flow around mountains, the difference is less than in  
807 the ocean where water must completely change direction.

808 Under the 1998 California standards, students discussed convection in both 5<sup>th</sup>  
809 and 6<sup>th</sup> grade, but under the CA NGSS this instructional segment is likely the first time  
810 students encounter the concept of convection. They will therefore need hands-on  
811 experience with the process in order to develop mental **models** of convection. These  
812 models begin with simple visualizations of convection using miso soup, rheoscopic fluid,  
813 or food coloring with water that allow students to recognize some general **patterns** of  
814 motion. They can then conduct more detailed **investigations** mapping out the motion of  
815 individual particles to provide evidence that supports the **argument** that uneven heating  
816 **causes** these patterns<sup>7</sup>. Students should be able to apply their model of convection to  
817 predicting the direction wind or water will move when exposed to uneven heating at the  
818 regional **scale** (a part of MS-ESS2-6). In India, changes in the heating differential  
819 between winter and summer cause the prevailing wind direction to reverse direction  
820 almost completely, creating their famous monsoons. Along the California coastline, we  
821 see this effect every day as the wind switches direction from morning to evening as the  
822 temperature difference between land and water switches direction.

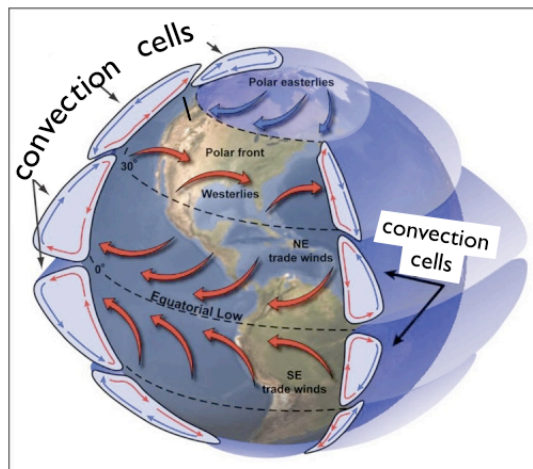
---

<sup>7</sup> UCAR, Atmospheric Processes-Convection: [https://www.ucar.edu/learn/1\\_1\\_2\\_7t.htm](https://www.ucar.edu/learn/1_1_2_7t.htm)

823           Understanding how convection works at the global **scale** helps explain many  
824 **patterns** in wind and precipitation. The strong temperature difference between equator  
825 and poles sets up convection, but as air masses move northward, some of their **energy**  
826 **flows** to their surroundings through cooling and drag. As a result, air from the equator  
827 does not make it all the way to the poles before it sinks back to the surface. Instead, our  
828 present-day atmosphere involves three major convection cells divided into latitudinal  
829 bands (**Figure 5**). Regions at the boundary between these convection cells tend to be  
830 areas with more dramatic weather: where both convection cells have air rising,  
831 thunderstorms are generated while the convergence between air masses at the upper  
832 mid-latitudes typically gives rise to rainier weather patterns.

833 Climate **patterns** are not permanent, and **changes** to the energy balance on the planet  
834 can **cause changes** to convection. The convection cells migrate with the seasons as  
835 well as local temperature variations. We see these changes as migrations of the jet  
836 streams, high velocity winds that race in the upper atmosphere along the boundary  
837 between convection cells. In the winter time, the convection cell boundary moves  
838 towards the south, bringing California its rainy winters that dry up in the summer as the  
839 convection boundary migrates back northward. Southern Europe is located at a similar  
840 latitude, so it has a similar pattern of weather, which is why our climate is often referred  
841 to as a “Mediterranean climate.” Students may not realize that large portions of the  
842 planet actually get the majority of their rain in the summer and that our unique climate is  
843 used to our unique position on the globe.

844           In addition to seasonal **changes** to the energy balance on the planet, **changes**  
845 at longer **timescales** can also occur. Computer simulations show that in periods of  
846 geologic history when there was a smaller temperature differential between the equator  
847 and poles, Earth may have had one large convection cell for each hemisphere spanning  
848 the entire region from equator to pole. Future climate changes may again disrupt wind  
849 and ocean currents.



850

851 **Figure 5.** Latitudinal bands in earth’s atmospheric circulation. (NC State University  
852 2013)

853

#### 854 **Additional background for teachers on Coriolis effects**

855 If simple convection were the only process controlling air movements, all wind  
856 would flow in the north-south direction, but we know that is not true. Earth’s rotation  
857 modifies this path. The assessment boundary for *MS-ESS2-6* states that “Assessment  
858 does not include the dynamics of the Coriolis effect,” so the exact details of this process  
859 are not essential for students but it may be desired by curious teachers and students.  
860 Air rotates around the Earth just like the planet overall. Material races around the  
861 equator at 1,700 km/hr to complete one full rotation in 24 hours, but it hardly needs to  
862 move at all near the poles. As a parcel of air travels from the fast moving equator  
863 towards the poles, it is moving faster in the direction of Earth’s rotation than the ground  
864 underneath it. From our perspective on the surface, it appears to be veering off in the  
865 direction of Earth’s rotation. Air moving from the poles towards the equator is moving  
866 slower than the ground underneath it, so it gets ‘left behind’ and appears to make a turn  
867 away from the rotation direction. Together, these deflections set up predictable bands of  
868 wind direction near the surface, and gives rise to the jet streams in the upper  
869 atmosphere.

## 870 **Angle of Sunlight and Seasons**

871 The angle of the Sun’s rays is also important for determining the variations in  
872 temperature during Earth’s seasons. Students combine their understanding of the effect  
873 of sunlight angle on energy input from this instructional segment with the orbital motions  
874 in the previous instructional segment to create a **model** that explains the reason for  
875 Earth’s repeating **pattern** of seasons (*MS-ESS1-1*). Students can make these  
876 connections using a physical model where their own body represents the motion of the  
877 planet<sup>8</sup>. They tilt their body towards or away from the Sun at the same 23.5° tilt as the  
878 Earth and move around Earth’s orbit, making sure that their tilt axis always points  
879 towards the North star. As they move from one side of the Sun to the other, they see  
880 how the angle of the Sun’s rays **changes** in the different hemispheres: in the northern  
881 hemisphere summer, the tilt brings the angle of the Sun’s rays closer to 90° while it  
882 makes the angle smaller in the southern hemisphere. Computer simulations allow  
883 students another way to visualize these changes<sup>9</sup>.

884 Learning a scientifically accurate model for the seasons is often impeded by  
885 students’ incoming preconceptions (documented vividly in the short documentary  
886 *Private Universe*<sup>10</sup> and in review articles<sup>11</sup>). Most notably, students often incorrectly  
887 believe that the Earth is closer to the Sun in summer and farther in winter. In this  
888 example course sequence, seasons are deliberately placed in a separate instructional  
889 segment from the discussion of orbits in order to increase the association between  
890 seasons and Sun angle instead of reinforcing an incorrect connection between seasons  
891 and orbital distance. Nonetheless, many students will still harbor this preconception and  
892 it must be addressed. Interactive 3-D simulations have been shown to help students

---

<sup>8</sup> Space Science Institute, Kinesthetic Astronomy.

[http://www.space-science-institute.org/education/extra/kinesthetic\\_astronomy/](http://www.space-science-institute.org/education/extra/kinesthetic_astronomy/)

<sup>9</sup> NOAA, Seasons and Ecliptic Simulator,

<https://www.climate.gov/teaching/resources/seasons-and-ecliptic-simulator>

<sup>10</sup> Harvard-Smithsonian Center for Astrophysics, Private Universe.

<http://www.learner.org/resources/series28.html?pop=yes&pid=9>

<sup>11</sup> <http://dx.doi.org/10.3847/AER2010035>

893 confront this preconception<sup>12</sup>. In these virtual worlds, students view the Sun-Moon-Earth  
894 **system** from various viewpoints and control different aspects, including rotation and  
895 revolution rates, and inclination of Earth’s spin axis. The story of seasons is mostly a  
896 story of light and energy absorption. Emphasis should be placed on the intensity and  
897 duration that sunlight shines on a particular patch of Earth’s surface. Because Earth’s tilt  
898 causes the Sun to appear to travel across the sky along a different path during summer  
899 versus winter, the Sun shines for longer days (causing longer duration sunlight) and  
900 from higher angles in the sky (causing more sunlight appear more intense in a given  
901 patch of the surface). Together, these give rise to warmer summers and cooler winters.

## 902 **Climate change**

903 Weather **changes** on many different **timescales**. There are trends and **patterns**  
904 that occur over hours, days, seasons, years, decades, and millennia. Shorter term  
905 variations are discussed in the next instructional segment. Scientists typically use the  
906 word ‘climate’ to describe patterns of weather that change over longer timescales. Many  
907 textbooks overemphasize the difference between the terms weather and climate; they  
908 are not different things but instead describe patterns and changes in atmospheric  
909 conditions over different timescales. The exact timescale that separates ‘weather  
910 patterns’ from ‘climate patterns’ is not universally agreed upon, but climate typically  
911 includes patterns that persist for decades or longer. Often, climate not only refers to the  
912 average conditions for a given location, but also includes a sense of the range of  
913 variation throughout the seasons and from year-to-year. Some climate changes may  
914 involve relatively small shifts to the average conditions, but substantially more frequent  
915 extreme weather (i.e., more severe droughts balanced by more extreme flooding or  
916 frequent heat waves balanced by frequent cold snaps).

917

---

<sup>12</sup> Something similar to this simulation is located here:  
[http://astro.unl.edu/naap/motion1/animations/seasons\\_ecliptic.html](http://astro.unl.edu/naap/motion1/animations/seasons_ecliptic.html), but it is described in  
Bakas and Mikropoulos 2003

918 **Common Core Connection**

919 Because temperature is a tangible topic and students have experience with its variation,  
920 climate data make an excellent way to engage students in 6<sup>th</sup> grade mathematics  
921 standards about statistics (*CA CCSSM.6.SP.A*).

922 **Changes** at each **timescale** are driven by different **causes**. Some climate  
923 changes in Earth's history were rapid shifts (caused by events, such as volcanic  
924 eruptions and meteoric impacts that suddenly put a large amount of particulate matter  
925 into the atmosphere or by abrupt changes in ocean currents). Other climate changes  
926 were gradual and longer term—due, for example, to solar output variations, shifts in the  
927 tilt of Earth's axis, or atmospheric change due to the rise of plants and other life forms  
928 that modified the atmosphere via photosynthesis. Scientists can infer these changes  
929 from geological evidence. Students can **analyze data** from these scientific observations  
930 to see how each process can correlate with observed changes in climate. Excellent data  
931 sets from tree rings and cherry blossoms exist showing how changes in sunspots and  
932 volcanic eruptions were recorded as changes in plant growth over the last 1,000  
933 years<sup>13</sup>.

934 Students begin to **analyze data** showing the temperature history over the last  
935 century (**Figure 6**). The focus in middle school is on **asking questions** about the  
936 **patterns** they see (*MS-ESS3-5*). In high school, students will build a **model** that can  
937 help explain the mechanisms causing the **changes** they see. While graphs like **Figure**  
938 **6** are simple enough for students to interpret, scientists also use more sophisticated  
939 interactive displays of data that depict how temperatures have changed in space and  
940 time. More advanced visualizations allow students to zoom into areas of interest (such  
941 as regions within California) and watch the time progression<sup>14</sup>. As students see the data  
942 depicted in new ways, they should be able to ask more detailed questions. For example,  
943 the right panel of **Figure 6** shows that the northern hemisphere has warmed more than

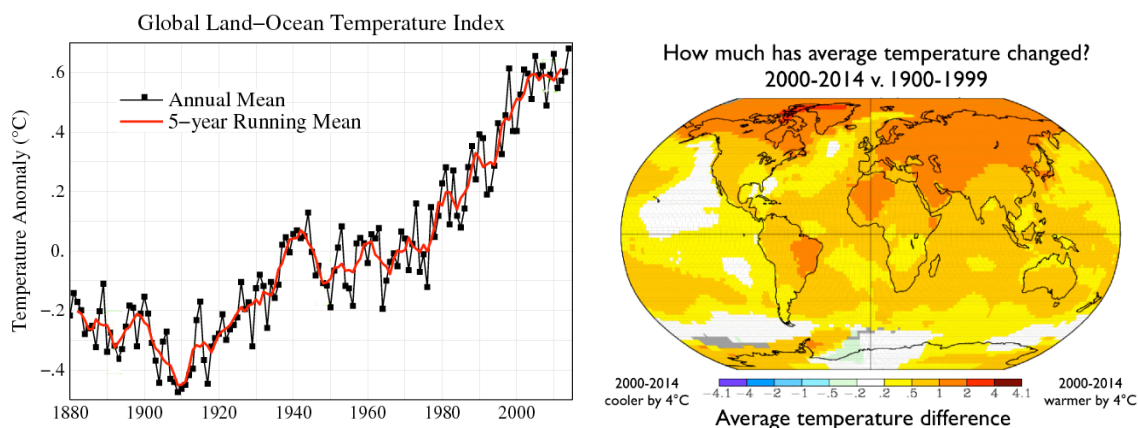
---

<sup>13</sup> National Center for Atmospheric Research, Investigating Climate Past: The Little Ice Age Case Study: <http://eo.ucar.edu/educators/ClimateDiscovery/LIA.htm>

<sup>14</sup> California Energy Commission, Cal-Adapt: <http://cal-adapt.org/tools/>

944 the southern hemisphere. Why? The eastern part of South America warmed more than  
 945 the west. Is that due to deforestation of the Amazon, or does it involve more complex  
 946 interactions? The lowest temperatures are shortly after 1900. What caused that? Did it  
 947 affect the whole planet equally? These are the types of **questions** we want our students  
 948 to start asking even though they won't have the tools to answer them yet in sixth grade.

949 The data also come alive when students **obtain information** about the **effect**  
 950 temperature **changes** have on sea-level, glaciers, or storm intensity. There are a  
 951 number of government reports summarizing these changes (EPA Climate Change  
 952 Indicators<sup>15</sup>, National Climate Assessment<sup>16</sup> or NASA's Climate Effects web portal<sup>17</sup>).  
 953 Students can research one aspect and prepare a summary product for the class that  
 954 **communicates** their findings.



955  
 956 **Figure 6.** Temperature changes over time depicted as a graph of average annual  
 957 temperatures for the entire globe since 1880 (left) and a map showing changes at  
 958 different locations, comparing the average from the first portion of the 21<sup>st</sup>  
 959 century to the 20<sup>th</sup> century (right). The 21<sup>st</sup> century is warmer than the 19<sup>th</sup> and 20<sup>th</sup>  
 960 centuries. (NASA 2015)

961

<sup>15</sup> EPA, Climate Change Indicators:

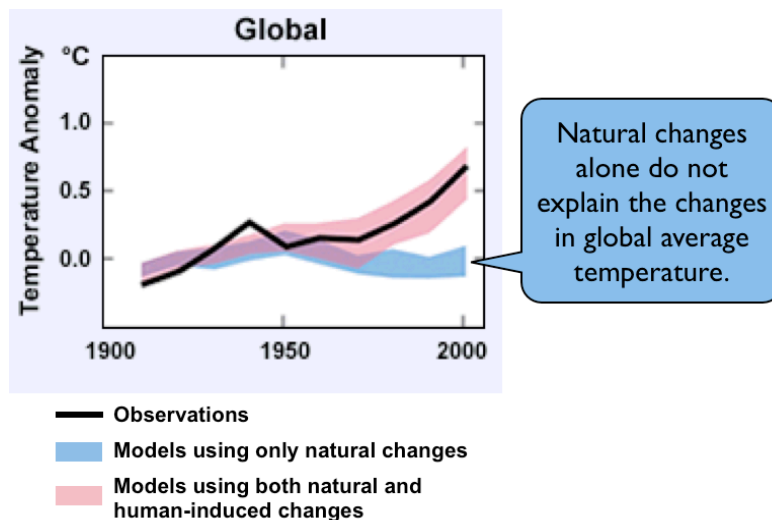
<http://www.epa.gov/climatechange/science/indicators/>

<sup>16</sup> National Climate Assessment: <http://nca2014.globalchange.gov/report#section-1946>

<sup>17</sup> <http://climate.nasa.gov/effects/>

962           There exist natural factors that can **cause** climate **changes** over human  
963 **timescales** (tens or hundreds of years), including variations in the Sun’s energy output,  
964 ocean circulation **patterns**, atmospheric composition, and volcanic activity (see  
965 ESS3.D). When ocean currents change their flow **patterns**, such as during El Niño  
966 Southern Oscillation conditions, some global regions become warmer or wetter and  
967 others become colder or drier. When scientists make computer simulations that include  
968 only these natural **changes**, they cannot match the temperature changes from the last  
969 century (**Figure 7**). But there are also changes that are caused by human activity  
970 (EP&C Principles III & IV). Many aspects of modern society result in the release of  
971 carbon dioxide and other greenhouse gases. These include automobiles, power plants  
972 or factories that use coal, oil, or gas as an energy source, cement production for  
973 buildings and roads, burning forest and agricultural land, and even the raising of  
974 livestock whose digestive processes emit methane. Greenhouse gases increase the  
975 capacity of Earth to retain energy, so changes in these gases cause changes in Earth’s  
976 average temperature. Changes in surface or atmospheric reflectivity change the amount  
977 of energy from the Sun that enters the planetary system. Icy surfaces, clouds, aerosols,  
978 and larger particles in the atmosphere, such as from volcanic ash, reflect sunlight and  
979 thereby decrease the amount of solar energy that can enter the weather/ climate  
980 system. Many surfaces that humans construct (e.g., roads, most buildings, agricultural  
981 fields versus natural forests) absorb sunlight and thus increase the **energy** in the  
982 **system**. As students **analyze data** about greenhouse gas concentrations in the  
983 atmosphere, they observe a very similar **pattern** to the change in temperature (**Figure**  
984 **8**). In fact, computer models of climate show that human activities are an important part  
985 of the **cause** of global temperature changes (**Figure 7**).





986

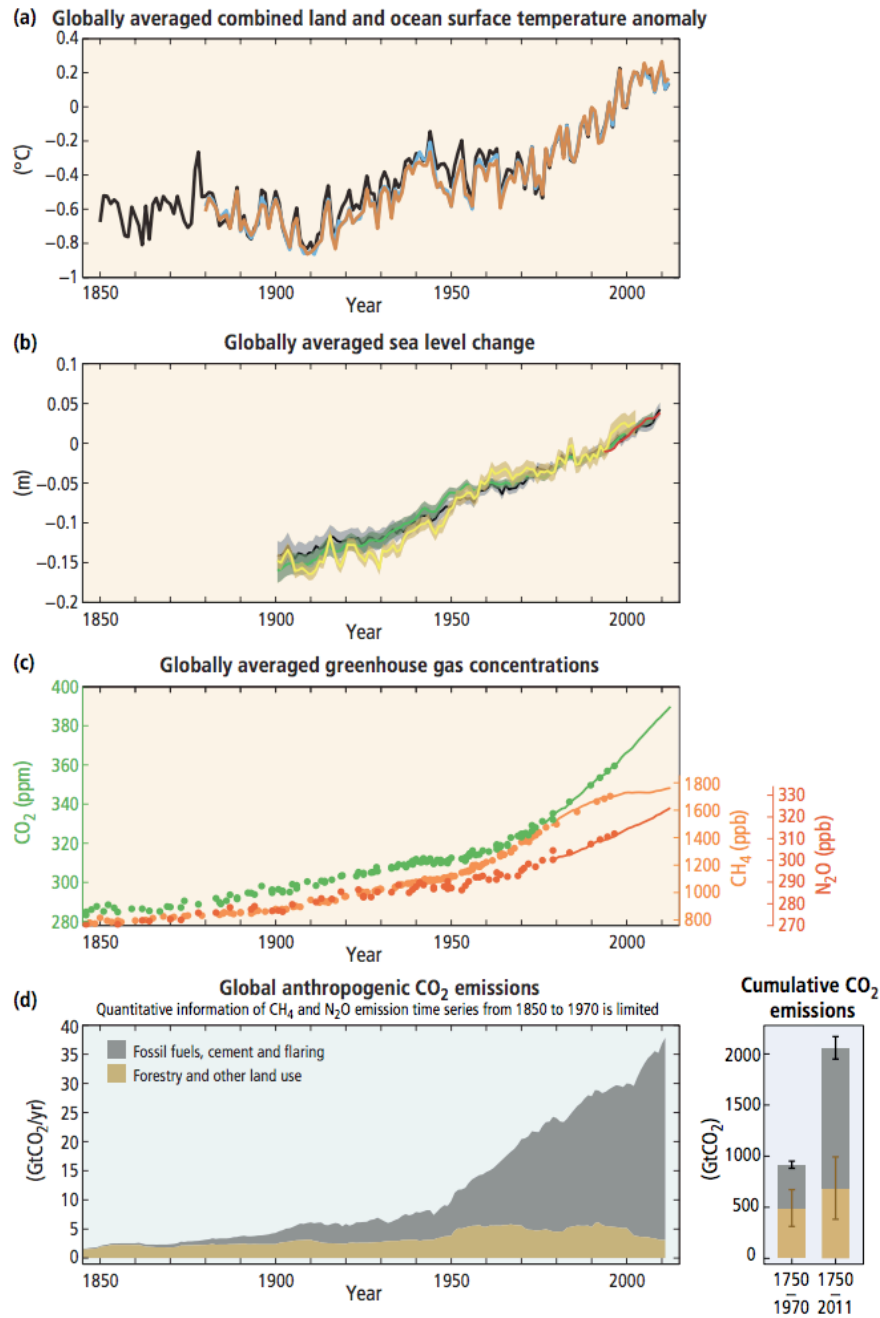
987 **Figure 7.** Outputs of different computer models of global climate compared to  
 988 observations. The colored bands are thick because they represent hundreds of different  
 989 models created by many different researchers using different assumptions. While the  
 990 models have slight variations in their output, only models that include human-induced  
 991 changes can explain the observed temperature record.

992

### 993 Common Core Connection

994 Global average temperature rises as humans emit more greenhouse gases. This rate of  
 995 emission depends on two key variables: population growth, and **energy** consumed per  
 996 person. Students must **construct an argument from evidence** that connects these  
 997 population and energy use ideas to a significant impact on Earth's systems (*MS-ESS3-*  
 998 *4*). To gather evidence for their argument, students **obtain information** from online  
 999 resources that list population and energy consumption **patterns**. Students will use  
 1000 **mathematical thinking** to create meaningful comparisons between the energy use in  
 1001 different states and countries. For example, energy use per person is an example of an  
 1002 'instructional segment rate' from ratio thinking in mathematics (*CA CCSSM 6.RP.2*).  
 1003 People in the US use more than twice as much energy per person than the average  
 1004 European country (U.S. Energy Information Administration 2015a), probably because  
 1005 our homes are bigger and spaced further apart. Californians, on average, use less  
 1006 energy per person than nearly every other state in the US (U.S. Energy Information  
 1007 Administration 2015b), partly due to our mild climate and partly due to effective energy

1008 efficiency programs. Despite this fact, the average Californian still uses more than 10  
1009 times more energy than the average person in the continent of Africa. These  
1010 comparisons are examples of ratios and ratio language (*CA CCSSM 6.RP.1*). Many  
1011 developing countries around the world have growing populations and are rapidly  
1012 changing their lifestyles to include more energy intensive tools. They will start  
1013 consuming energy at rates more like California or even the US average, which could  
1014 have a huge impact on global climate and global emissions. Computer **models** that  
1015 forecast **changes** in global climate rely on accurate estimates about energy  
1016 consumption in the future, and in high school students will use computer simulations to  
1017 explore the effects of these assumptions (*HS-ESS3-5*).



1018

1019 **Figure 8.** Graphs with similar trends and patterns illustrate global warming causes and  
 1020 effects. (Intergovernmental Panel on Climate Change 2014)

1021

1022 **Grade 6 Instructional segment 3: Atmosphere/Hydrosphere: Cycles of Matter**

Instructional segment 3: Atmosphere/Hydrosphere: Cycles of Matter	
Guiding Questions:	<ul style="list-style-type: none"> <li>• How do we predict tomorrow’s weather?</li> <li>• How do the atmosphere and hydrosphere interact to control our valuable water resources?</li> </ul>
Highlighted Scientific and Engineering Practices:	<ul style="list-style-type: none"> <li>• Developing and using models</li> <li>• Analyzing and interpreting data</li> <li>• Planning and carrying out investigations</li> </ul>
Highlighted Cross-cutting concepts:	<ul style="list-style-type: none"> <li>• <i>Energy and matter: Flows, cycles, and conservation</i></li> <li>• <i>Patterns</i></li> <li>• <i>Cause and effect</i></li> </ul>
Students who demonstrate understanding can:	<p><b>MS-ESS2-5. Collect data to provide evidence for how the motions and complex interactions of air masses results in changes in weather conditions.</b> [Clarification Statement: Emphasis is on how air masses flow from regions of high pressure to low pressure, causing weather (defined by temperature, pressure, humidity, precipitation, and wind) at a fixed location to change over time, and how sudden changes in weather can result when different air masses collide. Emphasis is on how weather can be predicted within probabilistic ranges. Examples of data can be provided to students (such as weather maps, diagrams, and visualizations) or obtained through laboratory experiments (such as with condensation).] [Assessment Boundary: Assessment does not include recalling the names of cloud types or weather symbols used on weather maps or the reported diagrams from weather stations.]</p> <p><b>MS-ESS2-6. Develop and use a model to describe how unequal heating and rotation of the Earth cause patterns of atmospheric and oceanic circulation that determine regional climates.</b> [Clarification Statement: Emphasis is on how patterns vary by latitude, altitude, and geographic land distribution. Emphasis of atmospheric circulation is on the sunlight-driven latitudinal banding, the Coriolis effect, and resulting prevailing winds; emphasis of ocean circulation is on the transfer of heat by the global ocean convection cycle, which is constrained by the Coriolis effect</p>

and the outlines of continents. Examples of models can be diagrams, maps and globes, or digital representations.] [Assessment Boundary: Assessment does not include the dynamics of the Coriolis effect.]

- MS-ESS3-2. Analyze and interpret data on natural hazards to forecast future catastrophic events and inform the development of technologies to mitigate their effects.** [Clarification Statement: Emphasis is on how some natural hazards, such as volcanic eruptions and severe weather, are preceded by phenomena that allow for reliable predictions, but others, such as earthquakes, occur suddenly and with no notice, and thus are not yet predictable. Examples of natural hazards can be taken from interior processes (such as earthquakes and volcanic eruptions), surface processes (such as mass wasting and tsunamis), or severe weather events (such as hurricanes, tornadoes, and floods). Examples of data can include the locations, magnitudes, and frequencies of the natural hazards. Examples of technologies can be global (such as satellite systems to monitor hurricanes or forest fires) or local (such as building basements in tornado-prone regions or reservoirs to mitigate droughts).]
- MS-ESS2-1. Develop a model to describe the cycling of Earth’s materials and the flow of energy that drives this process.** [Clarification Statement: Emphasis is on the processes of melting, crystallization, weathering, deformation, and sedimentation, which act together to form minerals and rocks through the cycling of Earth’s materials.] [Assessment Boundary: Assessment does not include the identification and naming of minerals.]
- MS-ESS2-4. Develop a model to describe the cycling of water through Earth’s systems driven by energy from the Sun and the force of gravity.** [Clarification Statement: Emphasis is on the ways water changes its state as it moves through the multiple pathways of the hydrologic cycle. Examples of models can be conceptual or physical.] [Assessment Boundary: A quantitative understanding of the latent heats of vaporization and fusion is not assessed.]
- MS-ESS3-4. Construct an argument supported by evidence for how increases in human population and per-capita consumption of natural resources impact Earth’s systems.** [Clarification Statement: Examples of evidence include grade-appropriate databases on human populations and the rates of consumption of food and natural resources (such as freshwater, mineral, and energy). Examples of impacts can include changes to the appearance, composition, and structure of Earth’s systems as well as the rates at which they change. The consequences of increases in human populations and consumption of natural resources are described by science, but science does not make the decisions for the actions society takes.]

Other necessary DCIs

PS3.B: Conservation of Energy and Energy Transfer

PS4.B: Electromagnetic radiation

Significant Connections to California’s Environmental Principles and Concepts:

Principle I. The continuation and health of individual human lives and of human communities and societies depend on the health of the natural systems that provide essential goods and ecosystem services.

Principle III. Natural systems proceed through cycles that humans depend upon, benefit from and can alter.

Principle IV. The exchange of matter between natural systems and human societies affects the long term functioning of both.

1023

## 1024 **Background and instructional Suggestions**

1025 California is known for its sunshine more often than its rain and snow, but it relies  
1026 on both of these to support its extremely productive agricultural sector and supply water  
1027 to its growing population (*EP&C I*). The previous instructional segment focused on  
1028 **energy flows** and briefly mentioned the **flow of matter** that enabled some of the  
1029 energy transfer. This instructional segment looks at the same processes from the  
1030 perspective of the **cycling of matter** in both the atmosphere and the hydrosphere  
1031 (*EP&C III*).

1032 The instructional segment on weather can be structured around the goal of  
1033 having each student create a weather forecast for their community. Classroom  
1034 instruction focuses on providing students the skills and background they need to  
1035 complete that task. The forecast theme allows students to explicitly name the  
1036 observable variables that describe their experience with weather: temperature, wind,  
1037 humidity, and precipitation, and air pressure. Even though the last variable, air pressure,  
1038 is crucial to understanding weather **changes**, an effective inquiry-based approach does  
1039 not introduce it as a key variable at the beginning. After all, people do not directly sense

1040 or feel changes in air pressure. Teachers focus on the observable quantities and then  
1041 encourage students to **ask questions** about what causes them to change.

1042         Students then **analyze data**, searching for **patterns** in the observable weather  
1043 variables that give clues about the **causes** of the **changes** (*MS-ESS2-5*). Students can  
1044 examine detailed maps of air pressure and wind patterns to discover that air moves  
1045 from high pressure to low pressure<sup>18</sup>. Understanding why this is true requires some  
1046 understanding of gases as particles. In 5<sup>th</sup> grade, students defined matter as particles  
1047 that are too small to see (*5-PS1-1*). In the discipline specific course sequence, they  
1048 have not yet **developed a model** of how those particles behave (it comes in 8<sup>th</sup> grade,  
1049 *MS-PS1-4*), so a partial model will need to be developed for this discussion. This model  
1050 simply defines high pressure as having lots of particles of air together in one place, all  
1051 moving and pushing against one another like people on a crowded dance floor. Lower  
1052 pressure regions are areas that have fewer particles packed together with more empty  
1053 space between them. Air particles from the crowded regions will get bumped and  
1054 pushed into those empty spaces such that there is an overall flow from high pressure to  
1055 low pressure.

1056         Students combine this **model** with their model of global convection from the  
1057 previous unit to create an even richer understanding of the movement of air and water  
1058 on Earth (*MS-ESS2-6*). The problem is best illustrated at a small **scale** along coastlines  
1059 where land masses are adjacent to water. Students conduct an **investigation** into the  
1060 thermal properties of land versus water to see how they heat and cool at different rates,  
1061 setting up temperature differentials. This uneven heating **causes** convection (the  
1062 movement of air) along coastlines just like it did on the global **scale** with the  
1063 temperature differential between the equator and the poles. As air heats up, the  
1064 particles spread out, with some of the warmer air rising upwards. The area where air  
1065 rises up is now lower pressure than its surroundings, so air begins to move from areas

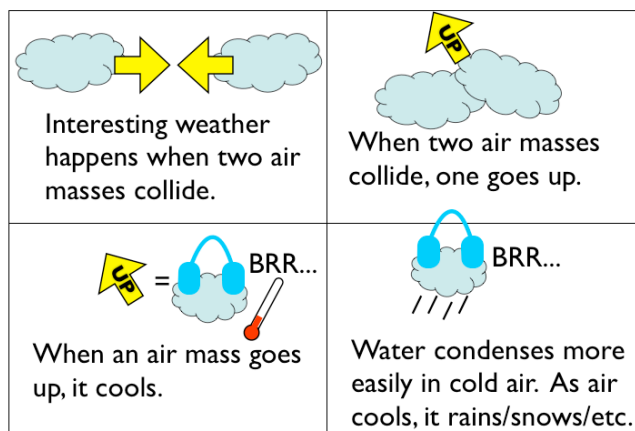
---

<sup>18</sup> American Meteorological Society, The Data Stream:  
<http://www.ametsoc.org/amsedu/dstreme/>

1066 where the pressure is higher (typically colder areas). Wind (the movement of air) results  
1067 from this convection cycle.

1068 The clarification statement for *MS-ESS2-5* indicates that students will not be  
1069 assessed on weather map symbols. This is largely a reaction to the fact that these  
1070 symbols are no longer necessary for illustrating weather **patterns** in the digital age. For  
1071 example, real-time wind patterns are indicated with animations of the flow of individual  
1072 particles<sup>19</sup> or with familiar rainbow color scales<sup>20</sup>. These visualization tools allow  
1073 teachers to spend more time helping students recognize and explain patterns with less  
1074 time devoted to memorizing symbols.

1075



1076

1077 **Figure 9.** Important components of a model of weather that describes the interaction of  
1078 air masses.

1079

1080 Using animations of real-time observations (such as satellite data from visible  
1081 light that reveals clouds and other wavelengths that reveal water vapor<sup>21</sup>), students  
1082 collect data about the movement of large air masses, noticing that the most intense

<sup>19</sup> Wind Map: <http://hint.fm/wind/>

<sup>20</sup> Earth: <http://earth.nullschool.net/#current/wind/surface/level/>

<sup>21</sup> NOAA, Geostationary Satellite Server: GOES Western U.S. Water Vapor:  
<http://www.goes.noaa.gov/browse3.html>



1083 precipitation and weather events occur where air masses collide (*MS-ESS2-5*). These  
1084 observations form the **evidence** that can be used to construct a complete **explanation**  
1085 or a **model** of the relationship between air masses and changing weather conditions.  
1086 The conceptual model in **Figure 9** shows that these explanations require further  
1087 investigation into condensation and the movement of water within Earth's systems.

1088 For a vignette related to weather, please see grade six of the Preferred  
1089 Integrated Model.

## 1090 **Water Cycle**

1091 At this point in Earth's history, very little water leaves the planet or arrives from  
1092 space. We simply need to track the movement of the matter that is already here. The  
1093 water cycle is therefore an example of a **cycle of matter** within a relatively closed  
1094 **system**. In 5<sup>th</sup> grade, students created graphs to illustrate where water is located on  
1095 Earth (*5-ESS2-2*) and they developed a **model** for the cycling of matter within the  
1096 biosphere (*5-LS2-1*). In 6<sup>th</sup> grade, they will extend their **model** to include the exchange  
1097 of water between all of Earth's **systems**, which should enable them to explain the  
1098 distribution of water they observed in 5<sup>th</sup> grade.

1099 Students hold many preconceptions about the way water is cycled through  
1100 Earth's systems (Ben-zvi-Assarf and Orion 2005). While they may be able to list the  
1101 locations where water can be found, they often are lacking a **model** for the  
1102 interconnectedness between these **systems** (i.e., water that is in the ground can flow  
1103 into rivers, oceans, or reach the surface at springs), or a sense for the dynamic  
1104 movement of water within each system (i.e., surface water doesn't just sit there waiting  
1105 to evaporate, but flows constantly down towards the oceans). Teachers can help  
1106 illustrate the dynamic interconnectedness of the water cycle through a simple  
1107 kinesthetic game. Students each play the role of a water molecule and will move around  
1108 the room through different stations that represent places where water is found on Earth  
1109 (ocean, lake, animal, plant, groundwater, atmosphere, ice cap, etc.). At each station,  
1110 they roll a dice and read from a table about the process that they will undergo so that  
1111 they can move from one station to another (i.e., evaporation, infiltration into the ground,

1112 flow downhill, come to the surface at a mountain spring, etc.). In essence, they become  
1113 a physical **model** for all the processes in the water cycle (*MS-ESS2-4*). The model  
1114 helps illustrate a number of concepts: 1) each of the reservoirs of groundwater are  
1115 interconnected; 2) water is constantly moving and flowing within each system and  
1116 between systems; 3) water is in different states (solid, liquid, and gas) in different  
1117 reservoirs, and **changes** in state (evaporation, condensation) are one key way that  
1118 water can move between different reservoirs; 4) there is no start, end, or single path  
1119 through the water cycle; and 5) changes in one part of the water cycle will have a major  
1120 impact on other parts of the system (e.g., if ice caps melt, sea level will rise; if the  
1121 climate warms and causes more evaporation, that will lead to more precipitation, which  
1122 will lead to more runoff).

1123 A model of the water cycle that only describes the **cycling of matter** does not  
1124 completely fulfill *MS-ESS2-4*, which requires students to explain how **energy**  
1125 exchanged via sunlight and gravity drives much of the movement. Additional  
1126 **investigations** into several of the process that cause movement of water through the  
1127 water cycle will help students understand these processes well enough to integrate  
1128 them into their **model**.

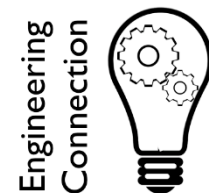
1129 **Energy** from sunlight has an **effect** on the water cycle because the increase in  
1130 thermal energy that it **causes** can in turn **cause** phase **changes**. Students conduct  
1131 **investigations** into evaporation and condensation to experience how they enable the  
1132 cycling of water and how they relate to **energy flow**. They recognize the **pattern** that  
1133 when water absorbs energy, it heats up and evaporates more readily. As water cools, it  
1134 tends to condense. Because they lack a detailed model of matter and state changes are  
1135 not introduced until 8<sup>th</sup> grade (*MS-PS1-4*), many students come away from these  
1136 activities believing the incorrect idea that water is the only material that can exist in all  
1137 three states of matter (after all, we only conduct this experiment with water and not  
1138 other materials). This preconception gets reinforced when students hear the true  
1139 statement that water is the only material that exists in all three states at the range of  
1140 natural temperatures on Earth (they seem to ignore the last part about natural  
1141 temperatures on Earth).

1142           The force of gravity also **causes** movement in the water cycle. Most students are  
1143 able to explain the role of gravity in precipitation (“raindrops fall”) or surface water  
1144 (“rivers flow downhill”), but often overlook the crucial role that gravity plays in infiltration  
1145 of surface water into the groundwater, the flow of groundwater itself through tiny pores  
1146 (illustrated as a saturated sponge drips water down out the bottom), and the flow of ice  
1147 downhill in glaciers (easily illustrated by time-lapse videos of glacier movement). In  
1148 order to emphasize these **cause and effect** relationships with gravity, students create  
1149 skits of different processes within the water cycle where one student is assigned to play  
1150 gravity or sunlight and must interact with other characters in the skit such as water  
1151 molecules or grains of sand. Short dramatic performances have been shown to improve  
1152 students’ conceptual understanding in science classes (Ødegaard 2003) and should  
1153 support language development. Drama comes in many forms, but can be particularly  
1154 well suited to **developing models of systems** by having individual characters play the  
1155 role of components within the system while their words and actions portray the  
1156 relationships between the components. The exchange of props between characters can  
1157 be a physical **model of cycles of matter**.

1158

### 1159 **Engineering solutions to pollution moved by the water cycle**

1160 Moving water often carries pollutants along with it (*EP&C IV*), but  
1161 understanding the water cycle allows people to design measures to  
1162 stop the flow of pollution. One possible engineering challenge for  
1163 students is to deal with the flow of water and pollutants in urban areas. As water runs  
1164 along road surfaces, it picks up oil, grit, and other pollutants that could flow into storm  
1165 drains and out into local waterways. During heavy rainstorms, those waterways can get  
1166 overloaded and flood. Allowing a greater fraction of water to infiltrate into the ground  
1167 can solve two problems because it reduces the amount of water on the surface that  
1168 causes flooding and the soil filters out many harmful contaminants before they flow  
1169 further. Students can be given the challenge of designing a system that divert waters



1170 into the ground and provides the maximum filtration of that water<sup>22</sup>. Students will have to  
1171 define specific criteria to measure their success (*MS-ETS1-1*), brainstorm and compare  
1172 different possibilities (*MS-ETS1-2*), test those possibilities (*MS-ETS1-3*), and make  
1173 iterative improvements (*MS-ETS1-4*).

1174

## 1175 **Human interaction with the Water Cycle**

1176 Because of the water cycle, Californians are able to obtain a steady supply of  
1177 fresh water for drinking, irrigation, industrial, and agricultural uses (*California EP&C III*).  
1178 Even in years with abundant precipitation, California still draws water from a total of  
1179 seven nearby states in addition to its own supply (The Nature Conservancy of California  
1180 2012). Of the water extracted for human use (“developed water”), more than 75% of it  
1181 goes to agriculture (California Department of Water Resources 2014) which helps  
1182 California grow more food than any other state (USDA 2015). While water is part of all  
1183 agriculture, some foods require more water to grow than others. If people choose to eat  
1184 more water efficient foods, California can cut back on its per-capita consumption of  
1185 water. Looking at data tables showing the water required for different food types,  
1186 students can compare the water footprint of several different meals. They will find that a  
1187 diet rich in meat products requires nearly twice as much water as a diet based on  
1188 vegetables and other plant products. For example, the average beef burger takes 4  
1189 times more water to produce than the same number of calories from an average soy  
1190 burger (Ercin, Aldaya, and Hoekstra 2012). The difference goes beyond water usage,  
1191 but includes other resources such as the land area required to grow the food and  
1192 **energy** resources to fertilize, transport, and process it. During their study of life science  
1193 in 7<sup>th</sup> grade and high school, students will learn more about food pyramids and the  
1194 concept of trophic levels that will help them understand why this should be the case. In  
1195 brief, predators inherently require more total energy input from the ecosystem than their  
1196 prey because of the energy used by the prey during its lifetime that is not preserved

---

<sup>22</sup> Engineering is Everywhere, Don't Runoff: Engineering An Urban Landscape:  
<http://www.eie.org/engineering-everywhere/curriculum-units/dont-runoff>

1197 within its biomass. Students can obtain global data about the relationship between  
1198 urbanization, rising incomes, and large increases in the amount of meat consumed per  
1199 person (expected to nearly double the levels from 1960 by the year 2030) (World Health  
1200 Organization 2015). With more people in the world eating more meat, there is  
1201 increasing pressure on water and other resources. Each family makes lifestyle choices  
1202 about the food they eat, and students should be able to construct an **argument** that  
1203 different lifestyle choices comes at the price of increased resource consumption (*MS-*  
1204 *ESS3-4*).

1205

1206

1207 **Grade 6 Instructional segment 4: Geosphere: Surface Processes**

Instructional segment 4: Geosphere: Surface Processes
<p>Guiding Questions:</p> <ul style="list-style-type: none"> <li>• How can we read layers of rock like the pages of a history book to reconstruct what happened during Earth’s past?</li> <li>• What is the relationship between the way rocks are built up (deposition) and the way rocks are broken down (erosion)?</li> <li>• How does our understanding of erosion and deposition help us find valuable energy and water resources and make ourselves safer from landslides?</li> </ul>
<p>Highlighted Scientific and Engineering Practices:</p> <ul style="list-style-type: none"> <li>• Constructing Explanations</li> <li>• Performing investigations</li> </ul>
<p>Highlighted Cross-cutting concepts:</p> <ul style="list-style-type: none"> <li>• <i>Structure and Function</i></li> </ul>
<p>Students who demonstrate understanding can:</p> <p><b>MS-ESS1-4. Construct a scientific explanation based on evidence from rock strata for how the geologic time scale is used to organize Earth’s 4.6-billion-year-old history.</b> [Clarification Statement: Emphasis is on how analyses of rock formations and the fossils they contain are used to establish relative ages of major events in Earth’s history. Examples of Earth’s major events could range from being very recent (such as the last Ice Age or the earliest fossils of homo sapiens) to very old (such as the formation of Earth or the earliest evidence of life). Examples can include the formation of mountain chains and ocean basins, the evolution or extinction of particular living organisms, or significant volcanic eruptions.] [Assessment Boundary: Assessment does not include recalling the names of specific periods or epochs and events within them.]</p> <p><b>MS-ESS2-1. Develop a model to describe the cycling of Earth’s materials and the flow of energy that drives this process.</b> [Clarification Statement: Emphasis is on the processes of melting, crystallization, weathering, deformation, and sedimentation, which act together to form minerals and rocks through the cycling of Earth’s materials.] [Assessment Boundary: Assessment does not include the identification and naming of minerals.]</p> <p><b>MS-ESS2-2. Construct an explanation based on evidence for how geoscience processes have changed Earth’s surface at varying time and spatial scales.</b> [Clarification Statement: Emphasis is on how processes change</p>

Earth’s surface at time and spatial scales that can be large (such as slow plate motions or the uplift of large mountain ranges) or small (such as rapid landslides or microscopic geochemical reactions), and how many geoscience processes (such as earthquakes, volcanoes, and meteor impacts) usually behave gradually but are punctuated by catastrophic events. Examples of geoscience processes include surface weathering and deposition by the movements of water, ice, and wind. Emphasis is on geoscience processes that shape local geographic features, where appropriate.]

**MS-ESS3-1. Construct a scientific explanation based on evidence for how the uneven distributions of Earth’s mineral, energy, and groundwater resources are the result of past and current geoscience processes.** [Clarification Statement: Emphasis is on how these resources are limited and typically non-renewable, and how their distributions are significantly changing as a result of removal by humans. Examples of uneven distributions of resources as a result of past processes include but are not limited to petroleum (locations of the burial of organic marine sediments and subsequent geologic traps), metal ores (locations of past volcanic and hydrothermal activity associated with subduction zones), and soil (locations of active weathering and/or deposition of rock).]

**MS-ESS3-2. Analyze and interpret data on natural hazards to forecast future catastrophic events and inform the development of technologies to mitigate their effects.** [Clarification Statement: Emphasis is on how some natural hazards, such as volcanic eruptions and severe weather, are preceded by phenomena that allow for reliable predictions, but others, such as earthquakes, occur suddenly and with no notice, and thus are not yet predictable. Examples of natural hazards can be taken from interior processes (such as earthquakes and volcanic eruptions), surface processes (such as mass wasting and tsunamis), or severe weather events (such as hurricanes, tornadoes, and floods). Examples of data can include the locations, magnitudes, and frequencies of the natural hazards. Examples of technologies can be global (such as satellite systems to monitor hurricanes or forest fires) or local (such as building basements in tornado-prone regions or reservoirs to mitigate droughts).]

Significant Connections to California’s Environmental Principles and Concepts:

Principle I. The continuation and health of individual human lives and of human communities and societies depend on the health of the natural systems that provide essential goods and ecosystem services.

Principle III. Natural systems proceed through cycles that humans depend upon, benefit from and can alter.

Principle IV. The exchange of matter between natural systems and human societies affects the long term functioning of both.

1208 **Background and instructional Suggestions**

1209           Every rock records a story. Earth scientists look out on a landscape and **ask**  
1210 **questions** about both the processes that are actively shaping it today and the specific  
1211 sequence of events in the past that led up to the present-day. Scientists **plan and**  
1212 **conduct investigations** to answer those questions, but investigations in Earth and  
1213 space science cannot always take the same experimental form with the testing of  
1214 hypotheses as they might in analytical chemistry or experimental physics. Many Earth  
1215 processes take millions of years and cover thousands of miles of area, time and  
1216 distance **scales** that are too slow and too large to reproduce in a lab. Geologists often  
1217 refer to the Earth as their ‘natural laboratory’, but they are only permitted to look at the  
1218 final result of its ancient experiments, or Earth’s present-day landscape. Investigations  
1219 in Earth science often begin with careful observations of what the Earth looks like today  
1220 and then try to reproduce similar features in small-scale laboratory experiments or  
1221 computer simulations.

1222           Students can develop this Earth science mindset when walking around their own  
1223 schoolyard and making observations about the familiar processes that led to its present-  
1224 day state<sup>23</sup>. For example, students can probably picture a brick wall being built layer-by-  
1225 layer, or that concrete starts off as grains of sand, gets mixed with cement, and then  
1226 hardens into solid ground. Not only can they observe those processes directly in their  
1227 everyday life, but they can see evidence of those processes as they walk around the  
1228 schoolyard. For example, looking closely at concrete, they can see different size grains  
1229 of sand held together with a grey material. As they look at these features, they realize  
1230 that they can **ask questions** about the world around them and how it came to look the  
1231 way that it does. Teachers can then introduce some natural geologic landscapes and

---

23

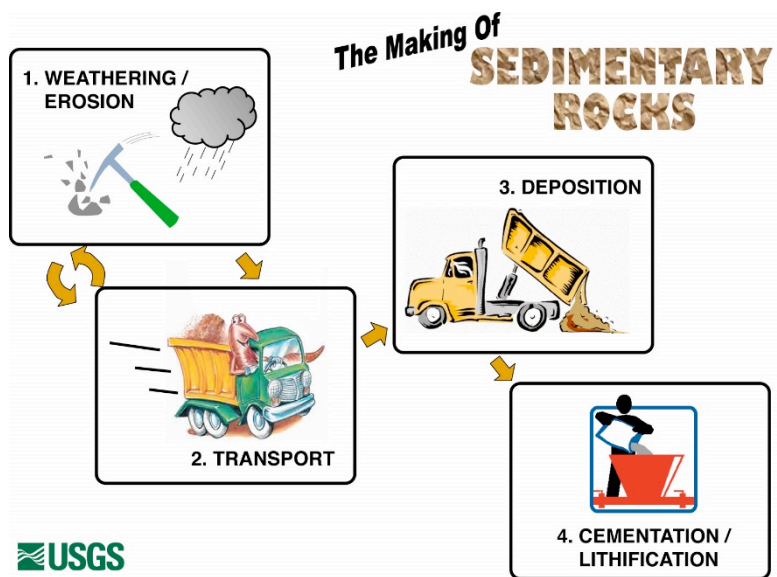
Schoolyard Geology, Lesson 3: <http://education.usgs.gov/lessons/schoolyard>



1232 processes that act on Earth and relate them to analogous processes from construction  
 1233 on the schoolyard.

1234 Earth scientists try to read layers of rocks like the pages of a history book. The  
 1235 composition and texture of each layer of rock reveals a snapshot of what the world  
 1236 looked like when that layer formed, and the sequence of layers reveals major events  
 1237 that reshaped them. In Earth science, these layers are the expression of the cross-  
 1238 cutting concept of **structure and function**. While in life sciences and engineering,  
 1239 structures are specific shapes so that they can accomplish certain functions, in Earth  
 1240 science structure is often a direct consequence of the processes shaping the planet.

1241 Each of these layers is built from material that came from somewhere else; this  
 1242 **cycle of matter** is referred to as the rock cycle. This instructional segment focuses on a  
 1243 portion of the rock cycle that occurs near the surface of the Earth where existing rocks  
 1244 are broken into pieces that are then moved around, reshaped, and combined back into  
 1245 a solid rock again. Rocks that are made directly from pieces of other rocks in these  
 1246 processes are called sedimentary rocks.



1247  
 1248 **Figure 10.** Processes involved in the making of sedimentary rocks. (USGS 2015)

1249

1250

1251 **Engineering connection: Cement and Sedimentary Rocks**

1252 Students may not realize it, but they are already familiar with  
1253 sedimentary rocks because most materials in the built  
1254 environment such as roads, sidewalks, bricks, and concrete are  
1255 essentially artificial sedimentary rocks with small pieces of rock material cemented  
1256 together. The average American is responsible for the use of nearly 9 tons of crushed  
1257 rock material every year of his or her life (USGS 1999b). These artificial materials are  
1258 carefully engineered to have sufficient strength at the lowest cost. Students can **obtain**  
1259 **information** about where rock aggregate comes from in their community (it is very  
1260 heavy and expensive to transport and usually quarried as locally as possible). The  
1261 process of cementation of natural sedimentary rocks usually occurs slowly underground  
1262 as mineral-rich water flows through pore spaces between grains, but it can be sped up  
1263 by adding concentrated cement minerals and water in a concrete truck. To develop a  
1264 **model** of how sedimentary rocks form (such as **Figure 10**; *MS-ESS2-1*), students can  
1265 engage in an engineering challenge to create the most durable concrete from plaster of  
1266 Paris and different size and shape rock pieces (sand, smooth pebbles, angular pebbles,  
1267 etc...) <sup>24</sup>. They decide the ideal **proportions** to mix the materials in small paper cups.  
1268 After letting their ‘concrete’ dry, they remove the paper cup and see whose material is  
1269 strongest by piling on different amounts of weight or dropping it from different heights  
1270 (*MS-ETS1-2*). This process helps motivate the rest of the instructional segment as it  
1271 provides students a physical model for the steps of sedimentary rock formation as well  
1272 as introducing them to the idea that rocks are broken down through the process of  
1273 erosion.

1274

1275 Students are now ready to apply their **model** for how sedimentary rocks form to  
1276 **constructing explanations** of how the Earth’s surface has **changed** over time (*MS-*  
1277 *ESS2-2*). The composition of the grains in a sedimentary rock matches the composition

---

<sup>24</sup> A short snippet for this idea is at:

<http://www.rsc.org/Education/Teachers/Resources/Inspirational/resources/4.3.2.pdf>

1278 of the original source rock. For example, a rock formation called the Gualala  
1279 conglomerate located near Point Arena in northern California contains large chunks of a  
1280 rock that appears to match the composition of the Gold Hill/Logan Gabbro in central  
1281 California. As rocks are transported by wind, water and gravity, the pieces are broken  
1282 down smaller and jagged edges are smoothed out over time. The Gualala conglomerate  
1283 has grains that are large, so they could not have traveled very far in a river before they  
1284 were deposited and cemented into a solid rock. In this case, however, some pieces of  
1285 the Gualala conglomerate are found hundreds of miles away from their matching source  
1286 rock. In addition to the small amount of movement by water and gravity, scientists infer  
1287 that these rocks were transported along half of California by the San Andreas fault over  
1288 millions of years! Students can perform an **investigation** similar to these scientists by  
1289 examining a sedimentary rock (in hand sample or photographs) and trying to match it to  
1290 different potential source rocks. They will have to **plan the investigation** by deciding  
1291 what features to observe in order to distinguish the different rocks and should consider  
1292 things like the size, shape, and composition of the grains. This form of investigation  
1293 where students are limited to observations and comparisons is common in Earth  
1294 science where it is often difficult to manipulate variables and experiments because the  
1295 time and spatial **scales** of Earth processes is so large.

1296         There are situations where Earth scientists can perform **investigations** that  
1297 simulate real-world processes at the small **scale**. A stream table (a sloped table or  
1298 plastic bin covered with sand and other earth materials and flooded with water) is a  
1299 platform for exploration about erosional processes and is an example of both a hands-  
1300 on **investigation** and a physical **model** that can be used to predict possible outcomes.  
1301 Students can use a stream table to investigate the factors that affect how quickly  
1302 material is broken off (weathered) and transported (eroded). When erosion is driven by  
1303 the movement of water, the steepness of the slope has a huge impact on the rate of  
1304 erosion because water builds up more kinetic **energy** when accelerating down a steep  
1305 hill (PS2.A). As the water molecules collide with the soil and rock, they can dislodge  
1306 individual pieces and carry them away. Students can also identify **patterns** in the  
1307 shapes of landforms in the stream table that might be similar to local landforms, such  
1308 steeply carved river channels that make meandering bends or wedge-shaped delta and

1309 alluvial fan deposits that form when the river reaches a flat section at the bottom of a  
1310 steep slope.

### 1311 **Forecasting Erosion Hazards: Landslides**

1312         Landslides are a rapid form of erosion that can damage property and put  
1313 peoples' lives at risk. Thankfully, areas that are most at risk for landslide hazard are  
1314 easy to recognize: steep slopes made of loose sediments are most at risk. Landslides  
1315 are also much more likely to happen during periods of intense rainfall, so their timing  
1316 can be forecast as well. Students can qualitatively explore these risk factors using a  
1317 stream table (a plastic tub filled with sand to represent Earth's surface and cups of  
1318 water as agents of erosion). They can perform **investigations** varying slope steepness  
1319 by changing the angle of the plastic tub, strength of different rocks by testing different  
1320 mixtures of clay and sand, and different rainfall intensities by using water containers  
1321 with different size holes. They can then **analyze data** from real landslides in their local  
1322 area using a state database of historical landslide studies<sup>25</sup> (*MS-ESS3-2*). Depending  
1323 on data availability in their area, their analysis could look for **patterns** in the sizes or  
1324 locations of landslides in comparison to the steepness of slopes or the types of rocks.  
1325 Because landslides tend to occur over and over again in the same regions, this type of  
1326 historical data helps inform the creation of maps of landslide hazard produced by the  
1327 state. Students can also **obtain information** from government agencies about efforts to  
1328 provide real-time forecasts of landslides in California so that people can either instigate  
1329 timely measures to reduce their hazard (these might include installing sandbags,  
1330 pumping water, or evacuating)<sup>26</sup>. This discussion has important ties to instructional  
1331 segment 3's discussion of weather patterns and the water cycle, but also to instructional  
1332 segment 1's discussion of climate **change**. In high school, students will explore how  
1333 landslide hazards could increase due to climate change (*HS-ESS3-5*).

---

<sup>25</sup> California Department of Conservation, Landslide and/or Liquefaction maps:  
<http://www.quake.ca.gov/gmaps/WH/regulatorymaps.htm>

<sup>26</sup> USGS, NOAA/USGS Demonstration Flash-Flood and Debris-Flow Early-Warning System: <http://landslides.usgs.gov/hazards/warningsys.php>

## 1334 **Depositing New Layers**

1335 Pieces of rocks and minerals, often called ‘grains’ when discussing rock  
1336 formation, get transported by gravity or moving wind and water. Conditions can **change**  
1337 such that there is no longer enough **energy** in these **systems** to continue to carry the  
1338 pieces, so they settle out and are deposited. For example, water moves fast as it races  
1339 down a steep hillside, but slows down when it reaches a lake or the ocean at the  
1340 bottom. As the water slows down, bigger grains settle first because they take the most  
1341 energy to move. In a classroom, the relationship between grain size and water velocity  
1342 can be illustrated in a large bottle filled with sand, soil, and water. After being shaken,  
1343 the largest grains of sand will fall out quickly, but the water at the top will remain muddy  
1344 for hours. When left overnight, the water will slow down enough that even the fine grains  
1345 will settle and leave clear, clean water at the top of the bottle. In nature, the deposition  
1346 of layers also buries any dead organisms and lead to the formation of fossils. By looking  
1347 at the types of organisms and the size of the grains, scientists can reconstruct the  
1348 geologic conditions in which the layer formed (i.e., was it a steep slope, where the river  
1349 meets the ocean, or far out to sea?). Sediment that is deposited later buries previously  
1350 deposited layers like a row of brick is placed on top of previously laid bricks to construct  
1351 a wall.

1352 Observing how layers **change** within a vertical sequence allows scientists to  
1353 track **changes** in the environment over time. The formation of mountain chains (that  
1354 push up mountains and therefore increase erosion) and ocean basins (new places  
1355 where rocks can settle out and be deposited) is gradual while volcanic eruptions, and  
1356 asteroid impacts are more abrupt. Periods of glaciation and warming occur at  
1357 intermediate timescales. **Changes** of all **timescales** are recorded as changes in the  
1358 rock layers and the fossils trapped within them. From this progression of layers,  
1359 geologists can reconstruct a timeline of the entire history of the Earth. Students likely  
1360 have heard of the names of geologic epochs like Jurassic, but exposure to these names  
1361 in middle school is distracting from the overall goal of using layers to determine the  
1362 relative timing of major events in Earth’s history. For example, major extinction events  
1363 are recorded in layers of rocks as decreases in the diversity of fossils around the world

1364 at the same period in geologic time. Students can **obtain information** from movies,  
1365 informational articles, or other resources in order to **construct an explanation** of how  
1366 evidence from layers of rock helped scientists identify a major event in geologic history  
1367 (*MS-ESS1-4*). Examples with a strong California focus include the extinction of the  
1368 dinosaurs 65 million years ago (a classic illustration of the nature of scientific discovery  
1369 that follows the work of University of California scientists)<sup>27</sup>, the eruption history of a  
1370 supervolcano like Long Valley caldera in eastern California, or the history of past glacial  
1371 periods determined by looking at layers in sediment cores taken from lakes in the Sierra  
1372 Nevada mountains.

### 1373 **Sediment Deposition, Groundwater Flow, and Energy Resources**

1374 The storage and flow of groundwater depends greatly on the materials that make  
1375 up the layers of rock and soil and how they formed. When layers of sediment are first  
1376 deposited, there is space between individual grains that water can flow through like  
1377 pores of a sponge. Sediment deposited in slowly moving water has small grains like silt  
1378 and mud with small and poorly interconnected pore spaces, so water does not flow well  
1379 through them. In environments where larger grains are deposited, the larger spaces  
1380 between grains tend to be well interconnected and water can flow through them easily.  
1381 Students can probably visualize dumping a bucket of water in a sandbox and having the  
1382 water flow quickly downward into the sand, but a muddy soil prevents the flow of water  
1383 and mud puddles can exist for hours after a rainstorm. A geologic setting where large  
1384 particles are deposited will lay down consistent layers of material that enables  
1385 groundwater flow. However, as the climate and environment **change** in cycles over  
1386 time, one location can alternate between layers of small grains and layers with larger  
1387 grains, interrupting the flow of water.

---

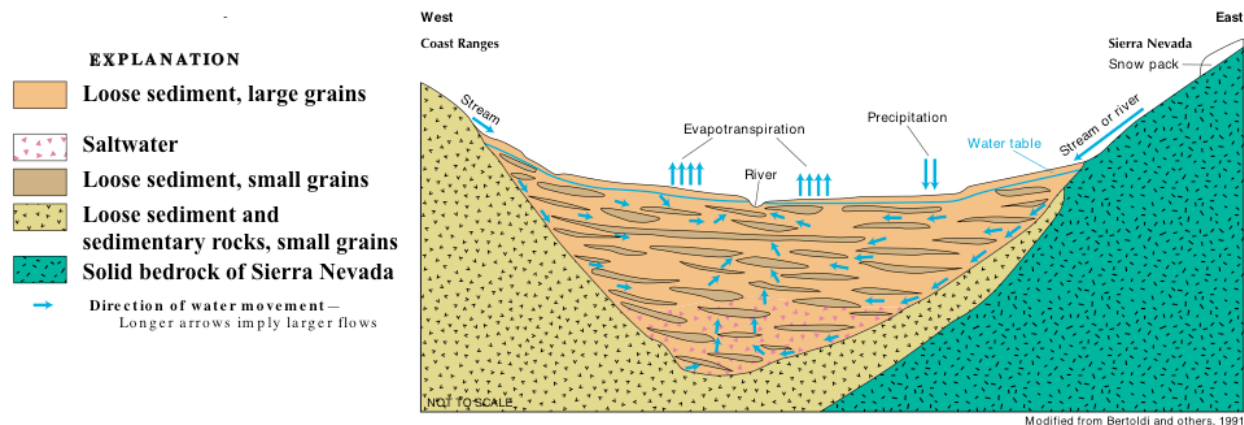
<sup>27</sup> Useful resources include: Howard Hughes Medical Institute, The Day the Mesozoic Died: <http://www.hhmi.org/biointeractive/day-mesozoic-died> and University of California Museum of Paleontology, Understanding Science: Asteroids and dinosaurs: Unexpected twists and an unfinished story: [http://undsci.berkeley.edu/article/0\\_0\\_0/alvarez\\_01](http://undsci.berkeley.edu/article/0_0_0/alvarez_01)

1388 Students can combine their **models** of sediment deposition and groundwater  
1389 flow to **construct an explanation** of how ancient geologic processes affect the present-  
1390 day distribution of groundwater resources in California (*MS-ESS3-1*). **Figure 11** shows  
1391 California’s Central Valley, which has accumulated thousands of feet of sediment that  
1392 eroded off the Sierra Nevada mountains during the last 100 million years. The  
1393 sediments are not all the same, however. There was once a shallow sea covering the  
1394 Central Valley, so only fine-grained sediments settled out to form layers. As plate  
1395 movements and climate **changed**, sea level changed and fast moving rivers flowing  
1396 over the land brought larger sized grains. As rivers changed their courses over time, the  
1397 size of grains being deposited at each individual location varied, leaving behind thin  
1398 lens-shaped layers of fine-grained sediments that impede the flow of water. In some  
1399 cases, these impermeable layers extend so far across the valley that they essentially  
1400 separate different pockets of groundwater from one another. These different pockets of  
1401 groundwater are a major source of water for farming in the Central Valley, especially in  
1402 years of drought when rain and snow do not provide sufficient surface water. While the  
1403 details differ, similar processes occur in other valleys of all sizes throughout the state.  
1404 When students look at a map showing the location of groundwater wells throughout the  
1405 state<sup>28</sup>, they should recognize the **pattern** that the vast majority of them are on valley  
1406 floors where layers of soft sediment have been recently deposited.

1407

---

<sup>28</sup> State Water Resources Control Board, Groundwater Ambient Monitoring and Assessment: <http://geotracker.waterboards.ca.gov/gama/>; USGS, National Water Information System Mapper: <http://maps.waterdata.usgs.gov/mapper/index.html>



1408

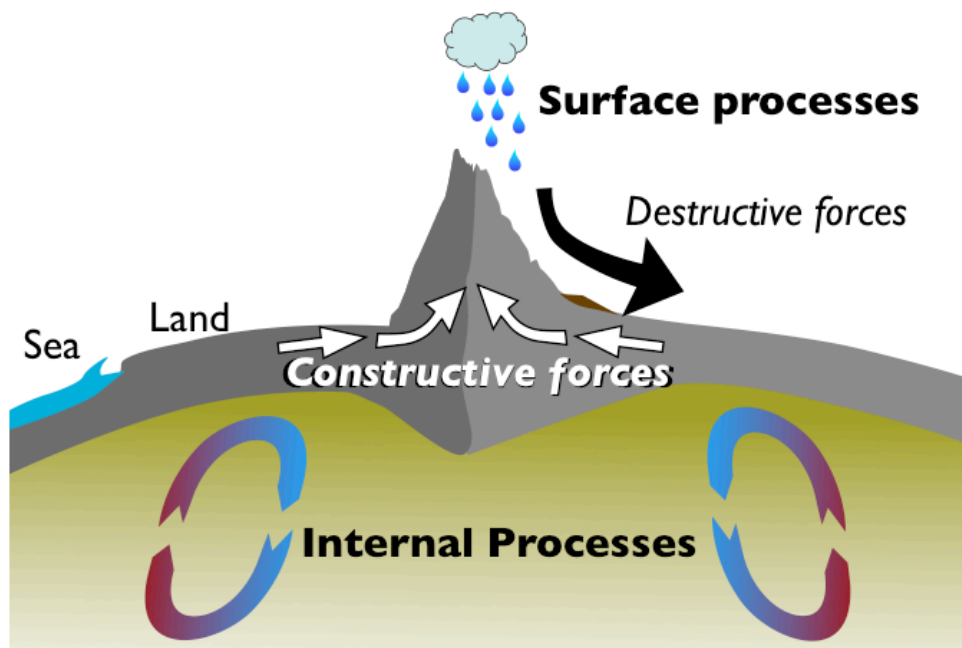
1409 **Figure 11.** A slice through California's Central Valley emphasizing groundwater flow  
 1410 through sediment in the ground as part of the water cycle. Water flows easily through  
 1411 the loose sediment made of large grains but does not penetrate easily into the small-  
 1412 grained loose sediments or the rocks beneath them. Gravity causes water to flow  
 1413 downward on both sides of the valley, but gets pushed back up when these flows  
 1414 converge along the axis of the valley, contributing water to the rivers and marshland in  
 1415 the Valley. (USGS 2009)

1416

1417 Water is not the only important resource that can flow through rocks; students  
 1418 can apply the same conceptual **models** to **explain** how crude oil and natural gas flow  
 1419 through pore spaces in rocks and can become trapped by layers with low permeability  
 1420 (*MS-ESS3-1*). Scientists working for oil and gas companies study the geologic history of  
 1421 an area so that they can target their drilling towards pockets where oil and gas is  
 1422 trapped. These scientists must also consider whether or not the geologic history of an  
 1423 area includes the deposition of large amounts of organic material (dead organisms)  
 1424 along with the original sediments. That organic material will slowly 'mature' into oil or  
 1425 gas resources through a series of chemical reactions sped up by the heat and pressure  
 1426 of burial. The reason these **energy** resources are so valuable is that it is rare to deposit  
 1427 layers in the ideal sequence for creating and preserving them: first a layer with  
 1428 abundant organic materials needs to be deposited, then a layer with large pore spaces  
 1429 through which oil and gas can flow and accumulate needs to be deposited on top of  
 1430 that, and then a layer with tiny grains to block the flow of oil and trap it at the right depth  
 1431 underground where it can be preserved for millions of years.



1432 In this instructional segment, students have focused solely on the development of layers  
1433 of sedimentary rock near Earth's surface and their relationship to the destructive force  
1434 of erosion. Many of the **changes** in what happens at the surface are in fact driven by  
1435 major changes inside Earth (**Figure 12**). The next instructional segment focuses on  
1436 those processes.



1437

1438 **Figure 12.** Landscapes are shaped at a range of timescales by processes inside the  
1439 Earth and on the surface. Image credit: (CC-BY-NC-SA) by M. d'Alessio

1440

1441

1442 **Grade 6 Instructional segment 5: Geosphere: Internal Processes**

Instructional segment 5: Geosphere: Internal Processes
<p>Guiding Questions:</p> <ul style="list-style-type: none"> <li>• How can the shapes of landforms at the surface help us understand processes that are going on deep within the Earth?</li> <li>• How can understanding plate motions help us locate resources (energy, mineral, and water) and protect ourselves from natural hazards?</li> </ul>
<p>Highlighted Scientific and Engineering Practices:</p> <ul style="list-style-type: none"> <li>• Analyze and Interpret Data</li> <li>• Develop and Use Models</li> </ul>
<p>Highlighted Cross-cutting concepts:</p> <ul style="list-style-type: none"> <li>• Patterns</li> <li>• Cycles of energy and matter</li> </ul>
<p>Students who demonstrate understanding can:</p> <p><b>MS-ESS2-1. Develop a model to describe the cycling of Earth’s materials and the flow of energy that drives this process.</b> [Clarification Statement: Emphasis is on the processes of melting, crystallization, weathering, deformation, and sedimentation, which act together to form minerals and rocks through the cycling of Earth’s materials.] [Assessment Boundary: Assessment does not include the identification and naming of minerals.]</p> <p><b>MS-ESS2-2. Construct an explanation based on evidence for how geoscience processes have changed Earth’s surface at varying time and spatial scales.</b> [Clarification Statement: Emphasis is on how processes change Earth’s surface at time and spatial scales that can be large (such as slow plate motions or the uplift of large mountain ranges) or small (such as rapid landslides or microscopic geochemical reactions), and how many geoscience processes (such as earthquakes, volcanoes, and meteor impacts) usually behave gradually but are punctuated by catastrophic events. Examples of geoscience processes include surface weathering and deposition by the movements of water, ice, and wind. Emphasis is on geoscience processes that shape local geographic features, where appropriate.]</p> <p><b>MS-ESS2-3. Analyze and interpret data on the distribution of fossils and rocks, continental shapes, and seafloor structures to provide evidence of the past plate motions.</b> [Clarification Statement: Examples of data include similarities of rock and fossil types on different continents, the</p>

	<p>shapes of the continents (including continental shelves), and the locations of ocean structures (such as ridges, fracture zones, and trenches).] [Assessment Boundary: Paleomagnetic anomalies in oceanic and continental crust are not assessed.]</p> <p><b>MS-ESS3-1. Construct a scientific explanation based on evidence for how the uneven distributions of Earth’s mineral, energy, and groundwater resources are the result of past and current geoscience processes.</b> [Clarification Statement: Emphasis is on how these resources are limited and typically non-renewable, and how their distributions are significantly changing as a result of removal by humans. Examples of uneven distributions of resources as a result of past processes include but are not limited to petroleum (locations of the burial of organic marine sediments and subsequent geologic traps), metal ores (locations of past volcanic and hydrothermal activity associated with subduction zones), and soil (locations of active weathering and/or deposition of rock).]</p> <p><b>MS-ESS3-2. Analyze and interpret data on natural hazards to forecast future catastrophic events and inform the development of technologies to mitigate their effects.</b> [Clarification Statement: Emphasis is on how some natural hazards, such as volcanic eruptions and severe weather, are preceded by phenomena that allow for reliable predictions, but others, such as earthquakes, occur suddenly and with no notice, and thus are not yet predictable. Examples of natural hazards can be taken from interior processes (such as earthquakes and volcanic eruptions), surface processes (such as mass wasting and tsunamis), or severe weather events (such as hurricanes, tornadoes, and floods). Examples of data can include the locations, magnitudes, and frequencies of the natural hazards. Examples of technologies can be global (such as satellite systems to monitor hurricanes or forest fires) or local (such as building basements in tornado-prone regions or reservoirs to mitigate droughts).]</p>
	<p>Significant Connections to California’s Environmental Principles and Concepts: none</p>

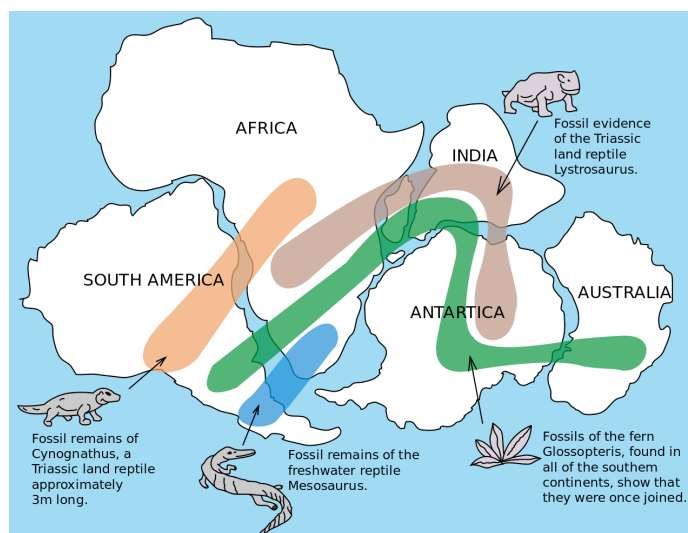
1443

1444 **Background and instructional Suggestions**

1445           If erosion were the only process sculpting Earth’s surface, all of the mountains  
 1446 would eventually wear away. While some of the mountains on Earth do look smooth and  
 1447 rounded because erosion has flattened them out, others look sharp and jagged as  
 1448 though they have not been exposed and weathered for very long at all. Is there a

1449 process that somehow renews mountain ranges, pushing them up so that erosion will  
 1450 then tear them down?

1451 In the early 1900's, a scientist named Alfred Wegener began looking at the  
 1452 locations of mountain ranges and noticed some **patterns**. He saw that the Appalachian  
 1453 mountains were made of the same unique rock types as the Scottish Highlands across  
 1454 the Atlantic, and that a mountain range in South Africa was similar to one in Brazil. He  
 1455 **asked questions** about what could possibly explain the large present-day separation,  
 1456 so he considered the idea that all of Earth's continents could have been connected  
 1457 together millions of years ago and subsequently moved to their current locations. He  
 1458 gathered substantial **evidence** that supported this proposed **explanation** and he began  
 1459 to refer to the idea as "continental drift."<sup>29</sup> Some of this evidence came from using maps  
 1460 to show how well the continents fit together, especially including the submerged  
 1461 continental shelves in aligning the continents, and most obviously with South America  
 1462 and Africa.



1463

1464 **Figure 13.** Fossil Evidence of Continental Drift. (USGS 1999a)

---

<sup>29</sup> An English translation of Wegener's 1912 article outlines the full range of his evidence: Wegener, A. 1912. *Die Entstehung der Kontinente* [The origin of continents] (Trans. R. von Huene), *Geol Rundsch* 3, 276-292.

[http://www0.unsl.edu.ar/~bibliogeo/index\\_archivos/wegener.pdf](http://www0.unsl.edu.ar/~bibliogeo/index_archivos/wegener.pdf)

1465

1466 Even more persuasive was **evidence** from fossils and rocks. **Figure 13** shows  
1467 continents from the Southern Hemisphere and how they could have been joined  
1468 together hundreds of millions of years ago. The colored areas correspond to fossils  
1469 whose specific geographic locations indicate not only that these continents were joined  
1470 together, but also specifically that the connection points match those predicted by  
1471 matching the outlines of the continents. The current wide separation of these continents  
1472 precludes other easy explanations for the locations of these fossils.

1473 Wegener also traced the past positions and motions of ancient glaciers based on  
1474 grooves cut by those glaciers in rocks, and also by rock deposits that the glaciers left on  
1475 different continents. His **evidence** indicated that if the continents had been in their  
1476 current locations, the glaciers would have formed very close to the equator, an  
1477 extremely unlikely situation. If the continents moved as he hypothesized, those glaciers  
1478 would have formed much closer to the South Pole.

1479 While we often say that Wegener compiled **evidence**, it is important to note that  
1480 he built on the work of dozens of scientists of the day. At the time Wegener lived, there  
1481 was no way to determine the exact age of rocks, but geologists could reconstruct the  
1482 relative timing of events by correlating sequences of rock layers from one place to  
1483 another (*MS-ESS1-4*, as discussed in instructional segment 4). Even though Wegener  
1484 never visited the Andes and the Atlantic coast of South America, other geologists had  
1485 written that folding of rock layers in the Andes mountain occurred at the same time as  
1486 rifting apart of the Atlantic ocean. Wegener **obtained and evaluated the information**  
1487 recorded by other scientists and then connected ideas in ways that nobody else had.

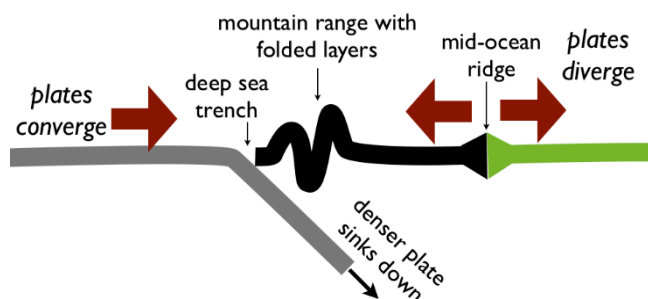
1488 Despite the **evidence** that he compiled, Wegener's theory was not accepted and  
1489 was generally forgotten. While Wegener was using traditional science practices of  
1490 **analyzing data** and **constructing explanations** based on **evidence**, the other  
1491 geologists were viewing his claims through the lens of the crosscutting concept of  
1492 "**cause and effect: mechanism and prediction.**" Wegener could not propose any  
1493 possible mechanism that would cause continents to plow through the ocean over great

1494 distances. In the absence of a mechanism to cause the proposed movements of  
1495 continents, the early twentieth century geologists rejected Wegener's claims. Middle  
1496 school students focus first on **analyzing the evidence** accumulated since Wegener's  
1497 time that provide even more definitive evidence that there has been motion of plates  
1498 (*MS-ESS2-3*), then they develop a **model** relating that motion to the **cycling of matter**  
1499 (*MS-ESS2-1*), and finally they can use that model to help explain **changes** in the  
1500 Earth's surface (*MS-ESS2-1*), the distribution of mineral resources (*MS-ESS3-1*), and to  
1501 forecast the occurrence of natural disasters (*MS-ESS3-2*). In high school, they will look  
1502 in more detail at some of the evidence and finally address the mechanism that drives all  
1503 this motion (*HS-ESS2-1, HS-ESS2-3*).

1504           Technological developments approximately 50 years later allowed detailed  
1505 mapping of the shape of the sea floor, which revealed new information that supported  
1506 Wegener's claims and also provided the missing mechanism. Students can investigate  
1507 undersea topography and notice **patterns** using a program like Google Earth. They can  
1508 discover that the largest mountain ranges on the planet actually exist below the water of  
1509 the ocean. One of the most obvious of these is the Mid-Atlantic Ridge, which rises about  
1510 3 km in height above the ocean floor and has a length of about 10,000 km running from  
1511 a few degrees south of the North Pole down almost all the way to the Antarctic circle.  
1512 While basically continuous across a huge part of the planet, it is far from straight. By  
1513 tracing out the shape of the continental shelves on either side of the Atlantic and the  
1514 axis of the Mid-Atlantic ridge, students can notice the ridge roughly parallels the turns of  
1515 the coastlines. By measuring the distance from the center of the mountain range to the  
1516 continental shelf, students can notice that the highest point of the mountains lies half  
1517 way between the two coastlines, as if the two coasts were spreading apart from this  
1518 central point. The idea that oceans were growing in size made it easier to understand  
1519 how the continents could move away from each other.

1520           With some ocean basins expanding, it didn't make sense for the entire planet to  
1521 be growing larger, so scientists began to look at how the growth could be balanced by  
1522 the surface appearing to get smaller in other locations. Scientists had long recognized  
1523 **evidence** for shortening on Earth because of evidence from sedimentary rock layers. In

1524 instructional segment 4, students created a **model** for how sedimentary rocks form in  
 1525 flat layers, but these layers are often observed to be folded and curved, which could only  
 1526 happen by some sort of squeezing that would push up mountains. At the time Wegener  
 1527 lived, the only process that scientists could conceive of that could cause such  
 1528 squeezing was the overall contraction of the Earth as it cooled after being formed long  
 1529 ago. If the seafloor was known to spread at some locations, it makes sense that plates  
 1530 must crash together at others. This would explain why mountain ranges formed long  
 1531 bands perpendicular to the spreading directions. For example, the Andes mountains are  
 1532 not oriented randomly – they are at exactly the orientation you would expect if South  
 1533 America was spreading away uniformly from the Mid-Atlantic Ridge and crashing into  
 1534 the Pacific ocean on the other side. Seafloor structures also give one more key piece of  
 1535 **evidence** about plate motions: there are very deep canyons in the ocean that parallel  
 1536 coastlines and island chains in many locations. Just off the west coast of South  
 1537 America, students can notice a very deep trench in the ocean floor. A physical **model**  
 1538 with two foam blocks (or even notebooks) representing plates helps illustrate why such  
 1539 a trench might form when one of the plates sinks down beneath the other. It is just a  
 1540 simple consequence of the geometry of a bending block, with the trench forming at the  
 1541 inflection point where the down going block starts to curve. Students can use maps of  
 1542 global topography and bathymetry to see if they notice any **patterns** between the  
 1543 location of these deep sea trenches and their relationship to continents, mountain  
 1544 ranges, and islands.



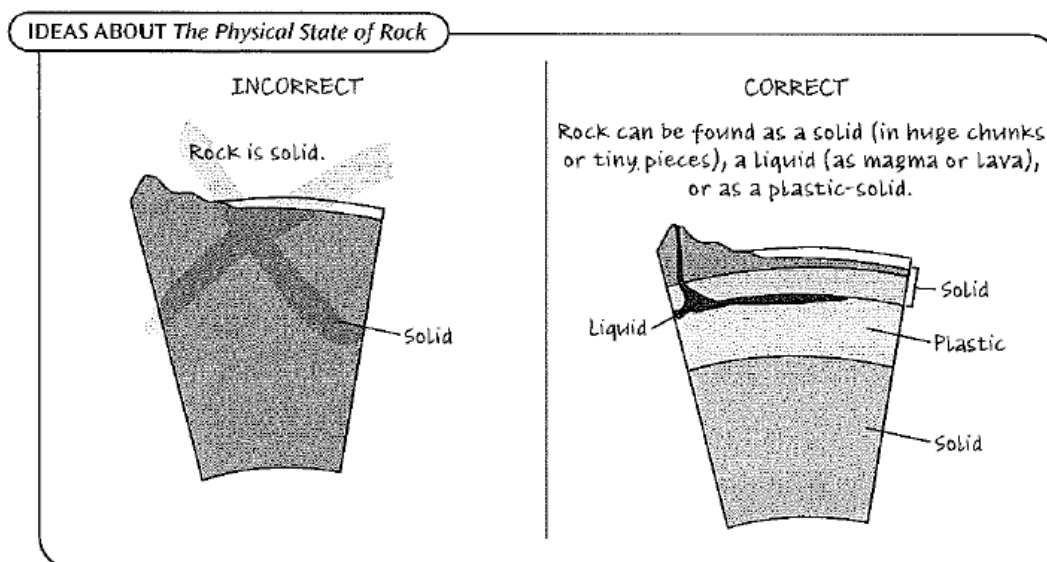
1545

1546 **Figure 14.** Schematic slice through the Earth's lithosphere showing three different  
 1547 plates with key seafloor and land features caused by their motion. Credit: (CC-BY-NC-  
 1548 SA) by Matthew d'Alessio

1549 Taken together, the fit of the continental shelves, the separation of similar rocks and  
 1550 fossils across vast oceans, the location of mid-ocean ridges running precisely along the  
 1551 center of oceans basins, the location of deep sea trenches along the coasts of some  
 1552 continents are strong **evidence** that plates move apart at some locations, together at  
 1553 others, and sliding past one another in other locations. These motions are the driving  
 1554 forces for a wide range of processes that shape earth's surface and cause interactions  
 1555 with the anthrosphere.

### 1556 **Plate tectonics drives the rock cycle**

1557 One of the most important effects of plate motions is the ***cycling of matter*** that  
 1558 accompanies the motion. The geoscience processes that form rocks and minerals  
 1559 include: volcanic eruptions, the heating and compaction of rock deep underground, the  
 1560 cooling of very hot underground rock, the evaporation of mineral-rich water, and the  
 1561 physical and chemical breakdown of surface rock by wind and water. All but the last of  
 1562 these geoscience processes are driven by the transfer of Earth's internal thermal  
 1563 **energy**. This internal thermal energy resulted from the immense heating of Earth's  
 1564 interior during its cataclysmic formation billions of years ago, the gravitational  
 1565 compaction of Earth in its early history, and the energy released by radioactive decay of  
 1566 buried Earth materials. In high school, students will develop a model that relates these  
 1567 heat transport processes to the driving motions of plate tectonics (*HS-ESS2-3*).



1568



1569 **Figure 15.** Ideas about the Physical State of Rock. Image Credit: Used by permission of  
1570 WestEd Making Sense of Science project.

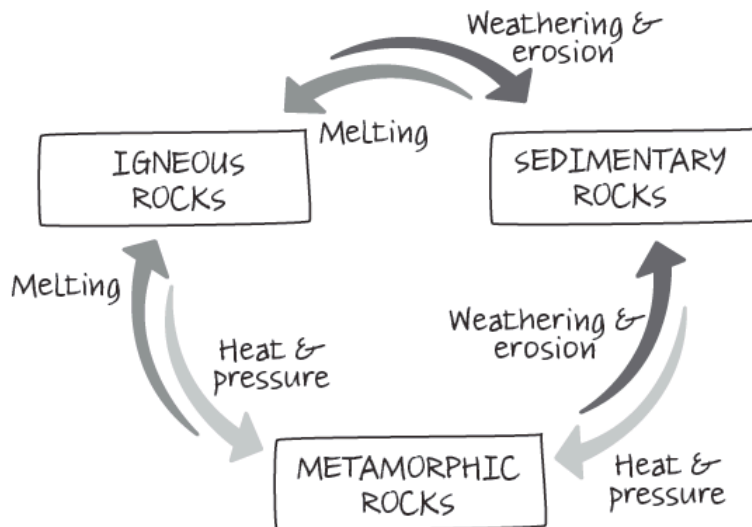
1571  
1572 Rock at Earth’s surface is almost exclusively a solid, except the few locations  
1573 where it flows as liquid lava. As shown in **Figure 15**, liquid rock is also located  
1574 underground, where it is called magma. Even in that illustration, the amount of liquid is  
1575 exaggerated for visual effect. A significant percentage of the rock underground exists in  
1576 a form that acts similar in some ways to a common children’s toy, silicone putty. It is not  
1577 clearly a solid or a liquid. This phase of matter is sometimes called a ‘plastic solid’  
1578 because it slowly flows and deforms under pressure like a liquid but retains its shape  
1579 like a solid. Even deeper underground, the immense pressure causes the rock to exist  
1580 as a solid.

1581  
1582 **Common Core Connection**

1583 Sometimes, everyday language differs from scientific language and can lead to  
1584 confusion. In **Figure 15**, the word ‘plastic’ refers to an easily shaped material. This  
1585 definition existed in dictionaries long before the invention of petroleum-based plastics  
1586 that we use so commonly in everyday materials like beverage bottles or bags. The  
1587 modern material called ‘plastic’ earned its name because it could be easily melted and  
1588 formed into different shapes. (CA CCSS for ELA/Literacy L.6.6)

1589  
1590 Many of the **changes** that happen to the geosphere (Earth’s nonliving solid  
1591 material excluding ice) are due to movement of tectonic plates. As the plates push  
1592 together, spread apart, and slide against one another, a variety of geologic processes  
1593 occur including earthquakes, volcanic activity, mountain building, seafloor spreading,  
1594 and subduction (sinking of a plate into the underlying mantle). All of these geoscience  
1595 processes change Earth’s rock – some form new rock, and others break down existing  
1596 rock.

1597



1598  
 1599 **Figure 16.** Classic Rock Cycle Diagram. Image credit: Used by permission of WestEd  
 1600 Making Sense of Science project.

1601  
 1602 These physical and chemical transformations of rock are often summarized as  
 1603 the rock cycle. **Figure 16** shows a classic rock cycle diagram with the three major rock  
 1604 types of igneous (melted in Earth’s interior), sedimentary (compacted from broken  
 1605 pieces), and metamorphic (rearranged by Earth’s internal pressure and thermal energy).

1606  
 1607 As summarized in

1608  
 1609 **Table 3**, the classic rock cycle diagram is a good summary of some of the key  
 1610 interactions of the geosphere. However, like most **models**, it has inaccuracies and can  
 1611 foster preconceptions. Students can mistakenly surmise that every rock has  
 1612 experienced or will experience the same cycle. However, rock does not move through  
 1613 the “rock cycle” in a specific order, like a product on a conveyor belt moving through a  
 1614 factory. The Geological Society in Britain has a very useful rock cycle website at  
 1615 <http://www.geolsoc.org.uk/ks3/gsl/education/resources/rockcycle.html>. This website is a  
 1616 very useful resource for students, who could then be challenged to find California  
 1617 examples of the British rocks and landforms.

1618

1619

1620

1621 **Table 3.** Benefits and Limitations of Classic Rock Cycle Diagram

Benefits	Limitations
Good summary of key geosphere interactions.	Does not show the many interactions the geosphere has with other Earth systems.
Easy to read and understand.	Does not show the timeframe for each geologic process, implying that they have similar timeframes.
Shows how each type of rock can become the other types of rock.	Does not show the locations where each geologic process takes place.
Helps dispel the incorrect idea that rock is “steady as a rock” and never changes.	Suggests that rock never leaves the rock cycle. Yet rocks often do leave the rock cycle, such as when they are incorporated into organisms, other Earth systems, and human-made materials.

1622 Table used by permission of WestEd Making Sense of Science project.

1623

1624 The physical and chemical **changes** that happen to minerals and rocks reinforce  
 1625 the principle of the **conservation of matter**. Almost three-quarters of Earth’s crust is  
 1626 made of oxygen and silicon. Just six elements (aluminum, iron, magnesium, calcium,  
 1627 sodium, and potassium) make up practically all the rest of Earth’s crust. Atoms of these  
 1628 eight elements combine to form Earth’s rocks and minerals. Throughout all the physical  
 1629 and chemical interactions, none of these atoms are lost or destroyed. Even as the  
 1630 appearance and behavior of the rocks **change**, their overall composition remains  
 1631 **stable**.

1632 Students can demonstrate that they understand the relationship between plate  
 1633 motion and the rock cycle by placing different types of rocks on an illustration showing  
 1634 typical plate boundaries (*MS-ESS2-1, MS-ESS2-2*). Magma solidifies to form igneous  
 1635 rocks at places where magma can reach the surface such as mid-ocean ridges. Rocks  
 1636 experience increases in temperature and pressure that can transform them into

1637 metamorphic rocks as they are dragged deep into the Earth when plates collide.  
1638 Sedimentary rocks form all over Earth's surface, but especially in zones where  
1639 mountains are actively being pushed up where plates collide.

## 1640 **Plate Tectonics and Resources**

1641 Plate tectonics plays an important role in the uneven distribution of Earth's  
1642 natural resources (*MS-ESS3-1*). Volcanic and uplift processes can bring important  
1643 minerals on or near the surface where they can be profitably mined. For example,  
1644 students can compare the location of the world's largest copper mines to the location of  
1645 plate boundaries and see that there is a general **pattern**: mines are often located near  
1646 plate boundaries. The prospector's shout that "there's gold in them thar hills" directly  
1647 connects gold distribution with the plate tectonics that created them thar hills.

1648 Fossil fuel distribution is one the most politically important uneven distributions of  
1649 natural resources, and it is also tied to plate tectonics. The Middle East has about 2/3 of  
1650 the world's proven reserves of crude oil. Petroleum and natural gas are generally  
1651 associated with sedimentary rocks. These fuels formed from soft-bodied sea organisms  
1652 whose remains sank to the ocean floor, decomposed in the relative absence of air, and  
1653 were further transformed by heat and pressure deep underground. Even areas on dry  
1654 land today can be the sites of ancient ocean basins that have been uplifted by plate  
1655 collisions. These same collisions can deform the rock layers in ways that allow oil and  
1656 gas to accumulate in concentrated locations (where they can be easily extracted) and  
1657 remain trapped there for millions of years. Students will **investigate** this process in high  
1658 school.

1659 Plate boundaries are often places where hotter material rises up from Earth's  
1660 interior to near the surface. This heat can be harnessed to generate electricity and as a  
1661 source of **energy** for heating buildings and commercial purposes. California is home to  
1662 some of the world's largest geothermal power plants, with production in both northern  
1663 and southern California that provide a total of 6% of the state's electricity (with potential  
1664 for even more). Other western states also utilize geothermal resources, but there are no  
1665 geothermal power plants east of North Dakota in the US, largely because these areas  
1666 are far from plate boundaries.

1667           In instructional segment 4, students learned about groundwater as an important  
1668 resource as water percolates into the spaces between pores in sediments and rock. The  
1669 distribution of groundwater basins is also affected by plate motions. The best  
1670 groundwater basins are in valleys where a large amount of sediment has continuously  
1671 been deposited, such as the Central Valley receiving sediment from the Sierra Nevada  
1672 mountains. Plate motions typically determine the shapes of these basins and are the  
1673 cause of mountains being uplifted in the first place. The faster they are pushed up, the  
1674 faster they erode (because rapid uplift produces steep slopes that erode quicker). Of  
1675 course, groundwater also requires an abundant source of water. In addition to the  
1676 important latitudinal controls on precipitation discussed in instructional segment 1,  
1677 mountains have a strong impact on where precipitation occurs; moist air flowing up  
1678 mountains tends to precipitate on the windward side of the mountains leaving a rain  
1679 shadow further downwind. The mountains that ‘squeeze moisture out’ are often recently  
1680 uplifted by plate motions.

#### 1681 **Understanding Plate Motions Allows Hazard Mitigation**

1682           In 4<sup>th</sup> grade, students analyzed **patterns** in maps and may have **investigated**  
1683 the distribution of earthquakes on the planet (4-ESS2-2). With an understanding of the  
1684 patterns of plate motions and previous events, scientists are better able to forecast  
1685 natural disasters such as earthquakes and volcanoes. The process is somewhat  
1686 analogous to asking students to predict where in California it will snow next January.  
1687 With a basic understanding of the patterns of geography, they could very reliably  
1688 identify places where it will almost certainly not snow (downtown Los Angeles, for  
1689 example) and where it is more likely to snow (perhaps along the high peaks in the  
1690 Sierra Nevada mountains). Whether it actually snows during that month depends on  
1691 specific physical processes, such as the location of the jet stream, which are difficult or  
1692 impossible to predict far in advance. Earthquakes occur because friction causes plates  
1693 to stick together where they touch. Even though forces deep within the Earth try to pull  
1694 them along, the plates remain stuck until the strain builds up so much that they  
1695 suddenly slide past one another in a single violent lurch. Students can build a physical

1696 **model** of this process with a brick, a bungee cord, and sand paper<sup>30</sup>, or explore a virtual  
1697 physical model using an online simulator<sup>31</sup>. Scientists can monitor the amount of strain  
1698 built up along plate boundaries using high precision GPS and can calculate the amount  
1699 of strain that is likely to be released in the next large earthquake at different locations. In  
1700 other words, scientists can predict where earthquakes could be and how big they could  
1701 be with relatively high reliability. State and local authorities have published maps  
1702 showing the likelihood of different size earthquakes in locations throughout California<sup>32</sup>.  
1703 Students could use this map to hold a mock session of the state legislature debating the  
1704 allocation of earthquake preparedness funding. Different students representing different  
1705 districts around the state use information about their population, their economic  
1706 contributions, and the earthquake forecasts to **argue** that their district is deserving of a  
1707 larger share of the funding.

1708         The only part of the process that is not yet predictable is the exact timing of the  
1709 earthquakes. While scientists have **investigated** a wide range of monitoring strategies,  
1710 it appears that many earthquakes occur without any perceivable trigger. That means  
1711 that the soonest we can know about earthquakes is the moment that they first start.  
1712 Earthquake waves do take time to travel through the Earth, so there is one more way  
1713 that understanding earthquakes can help us mitigate their effects. The moment a  
1714 seismic recording station detects shaking, it can send a signal at the speed of light to a  
1715 central processing center that can issue a warning of the impending earthquake. Such  
1716 warnings can be distributed to schools, businesses, and individuals via the internet,  
1717 mobile phones, and other broadcast systems, providing them warning of a few seconds  
1718 to a minute. Such systems have been in successful operation in Japan and Mexico City,  
1719 and a prototype is being tested in California. After investigating **patterns** of earthquake  
1720 occurrence in their region, students can make decisions about where to place seismic  
1721 recording devices to design their own earthquake early warning network that provides

---

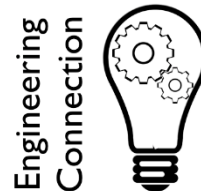
<sup>30</sup> SERC, Earthquake Demonstration:

<http://serc.carleton.edu/introgeo/demonstrations/examples/earthquake.html>

<sup>31</sup> CSUN, The Earthquake Machine: <http://www.csun.edu/quake/eqmachine>

<sup>32</sup> USGS. UCERF3: A New Earthquake Forecast for California's Complex Fault System, USGS Fact Sheet 2015-3009: <http://pubs.usgs.gov/fs/2015/3009/pdf/fs2015-3009.pdf>

1722 the maximum advance warning (*MS-ESS3-2*) (d'Alessio and Horey 2013). Using an  
1723 online simulator<sup>33</sup>, students test their network's performance in  
1724 sample earthquakes, compare network designs with their peers  
1725 (*MS-ETS1-2*) and iteratively improve them (*MS-ETS1-3*).



1727 **Common Core Connection**

1728 Students can use simple equations of distance, speed, and time to calculate the amount  
1729 of warning they can expect when a seismic recording station is a given distance away  
1730 from the earthquake source (*CA CCSSM 6.EE.2.c, 6.EE.7*).

---

<sup>33</sup> CSUN, Earthquake Early Warning Simulator: <http://www.csun.edu/quake>